

# In-Situ Stratospheric Size Distribution Measurements over the last 10 years – comparisons with OSIRIS and OMPS extinction, and COBALD backscatter measurements.

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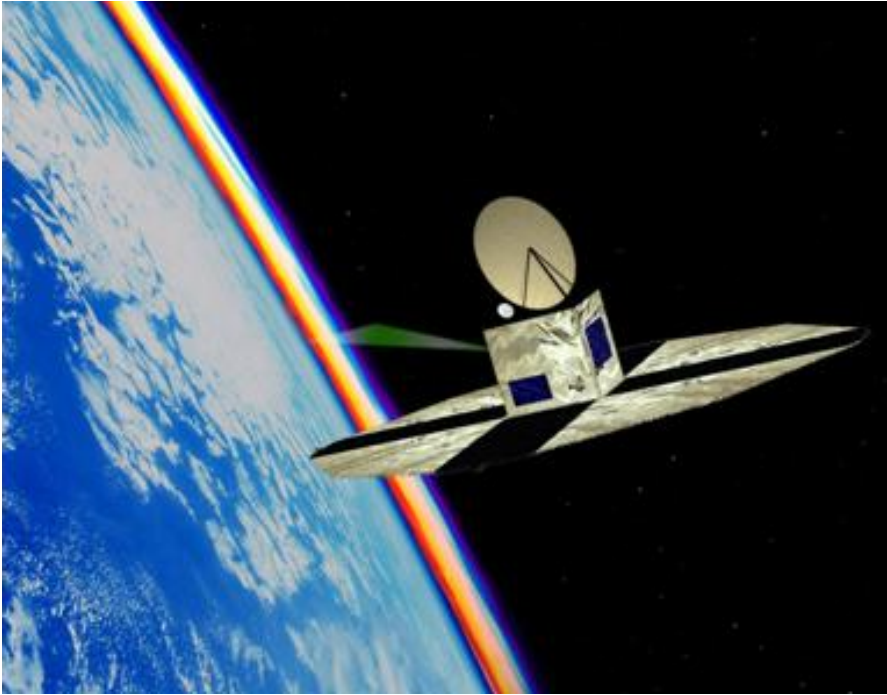
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<sup>4</sup>Swiss Federal Institute of Technology (ETHZ), Zurich, Switzerland

# University of Wyoming Optical Particle Counters (OPCs) – continuous measurements of stratospheric aerosol since 1971







Extinction retrieval from limb scattering

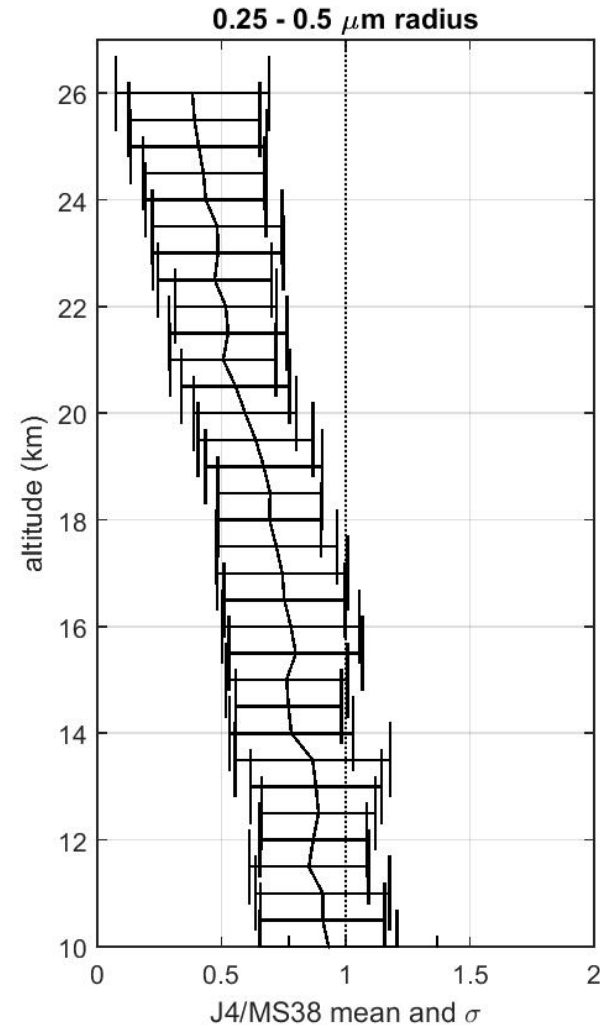
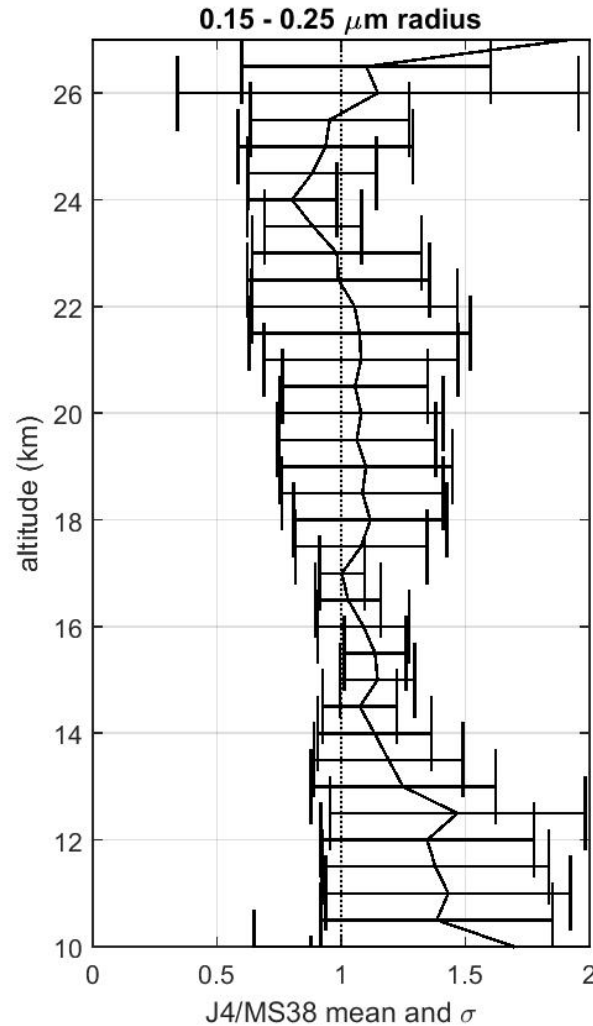


$N(r)$  – Concentration as function of radius

# Newest OPC compared to previous generation

10 coincident flights between 2006 – 2013 to characterize performance of latest OPC

$$\text{new OPC/old OPC} = J4/M_{\lambda}$$

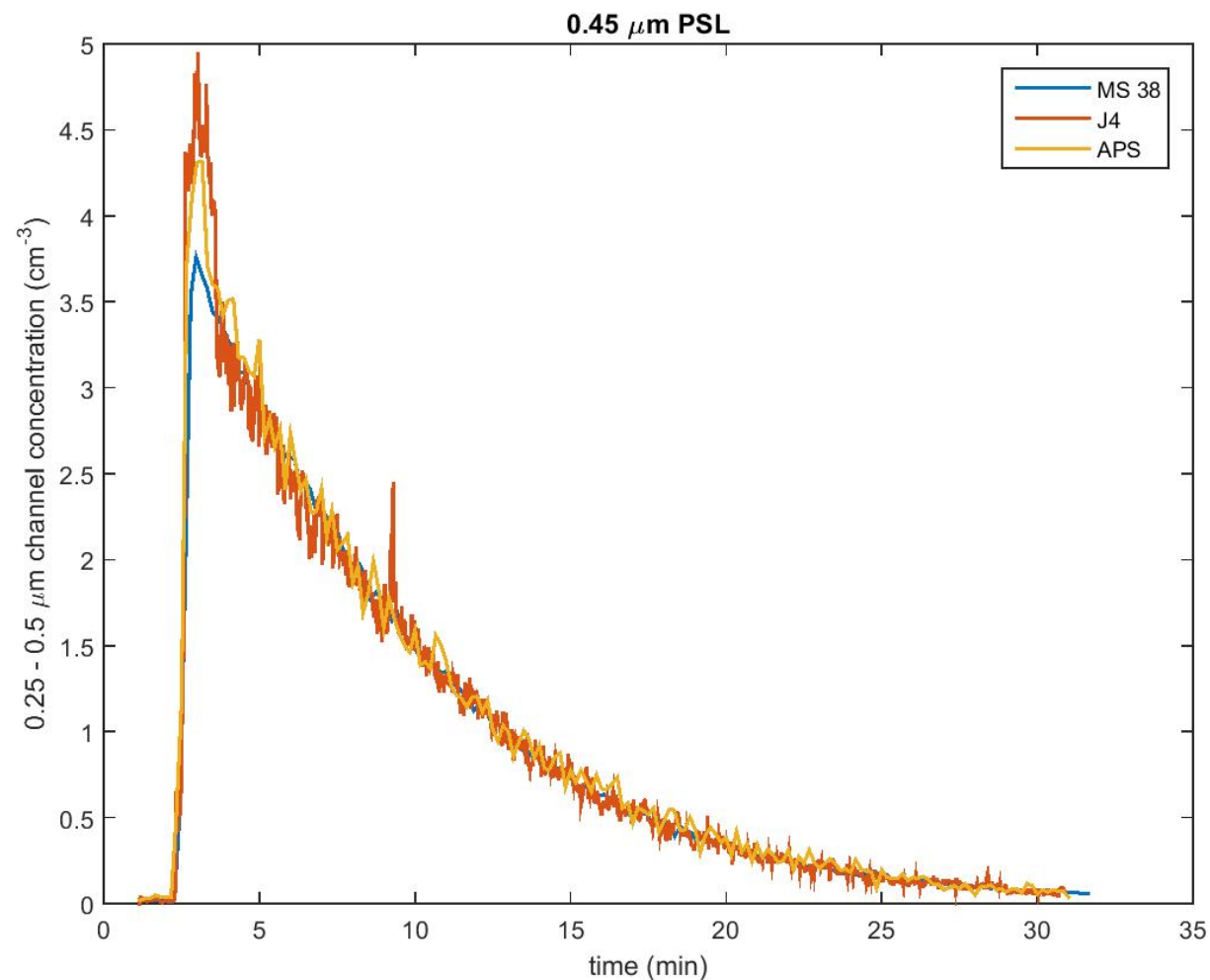


# Comparisons between 2 different UW OPCs

0.45  $\mu\text{m}$  radius PSL

This is upper end of the channel with a discrepancy  
(0.25 – 0.5  $\mu\text{m}$ )

Aerodynamic Particle Sizer as ground truth



# From size distributions to extinction

1. OPC measures  $N(r)$  in 8 size bins
2. Bimodal lognormal size distribution is fit to  $N(r)$  (Deshler, 2003)

$$N(>r) = \sum_{i=1}^8 \int_r^{\infty} N_i / \sqrt{2\pi \ln \sigma_i^2} \exp\left(-\ln^2[a/r_i] / 2 \ln^2 \sigma_i\right) d \ln a$$

3. Apply Mie theory to calculate extinction coefficient ( $Q$ ) as a function of particle radius, wavelength, and index of refraction

$$m = 1.45 - 0i$$

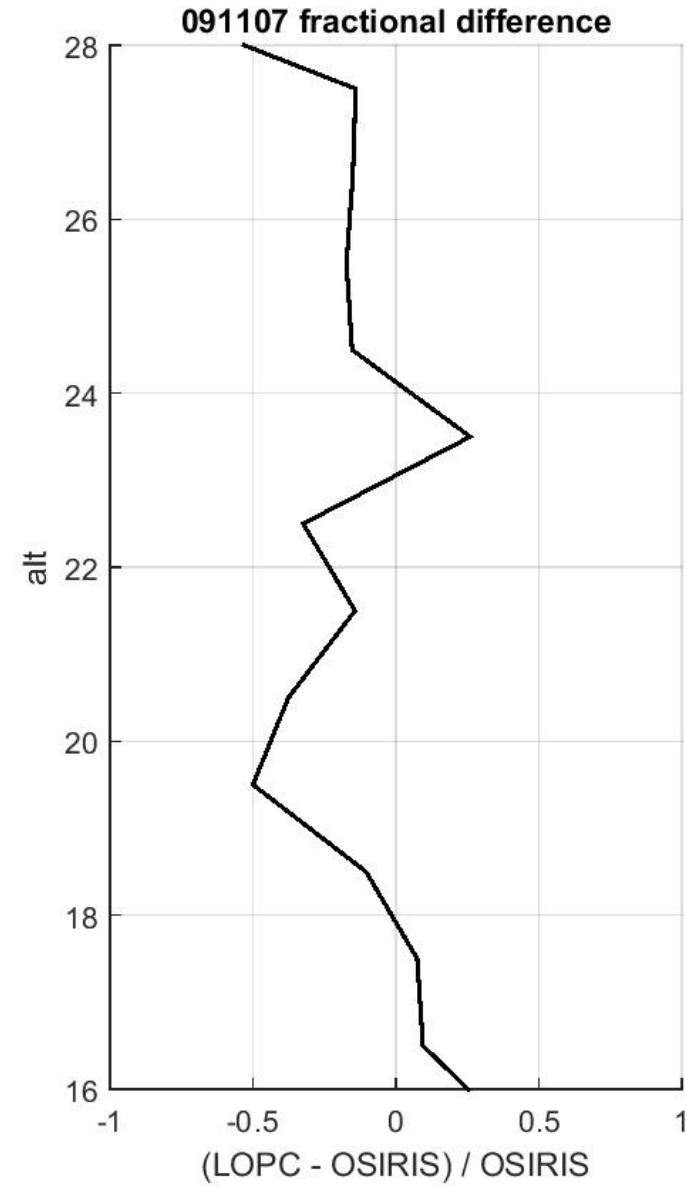
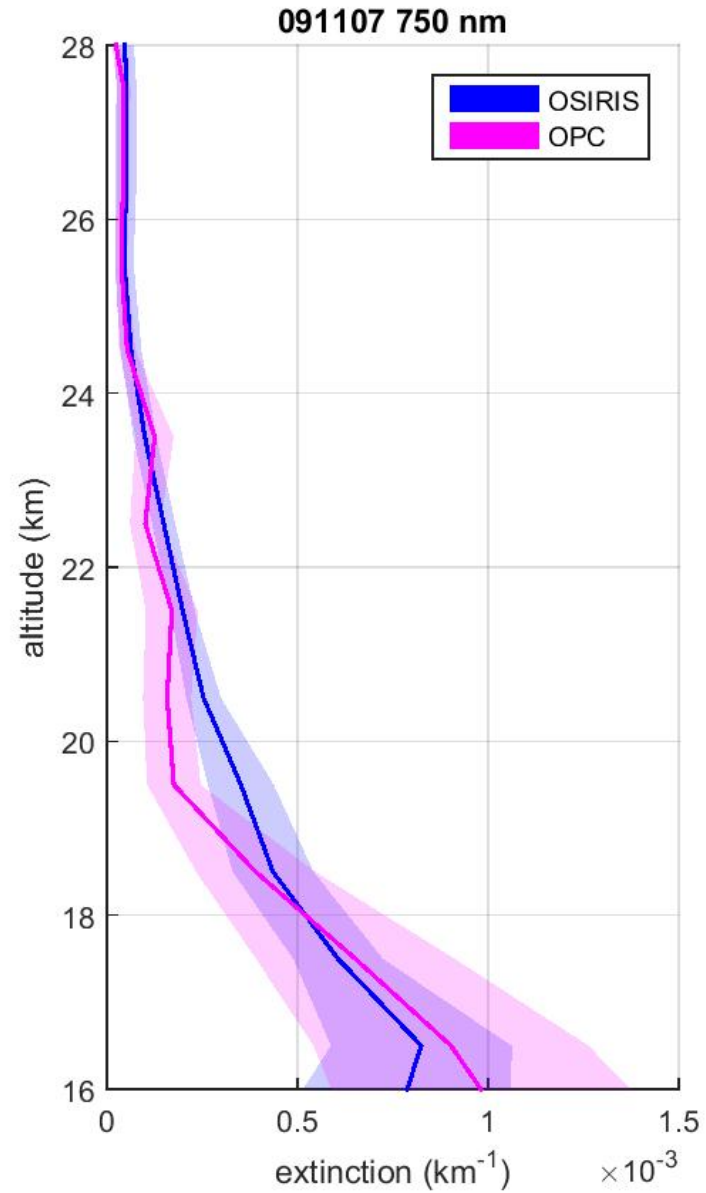
$$\beta_{\lambda} = \int_0^{\infty} \pi r^2 Q(r, \lambda, m) dn(r) / dr dr$$

# Comparisons to OSIRIS: 8 flights from 2006 - 2013

Large variability in extinction during this time



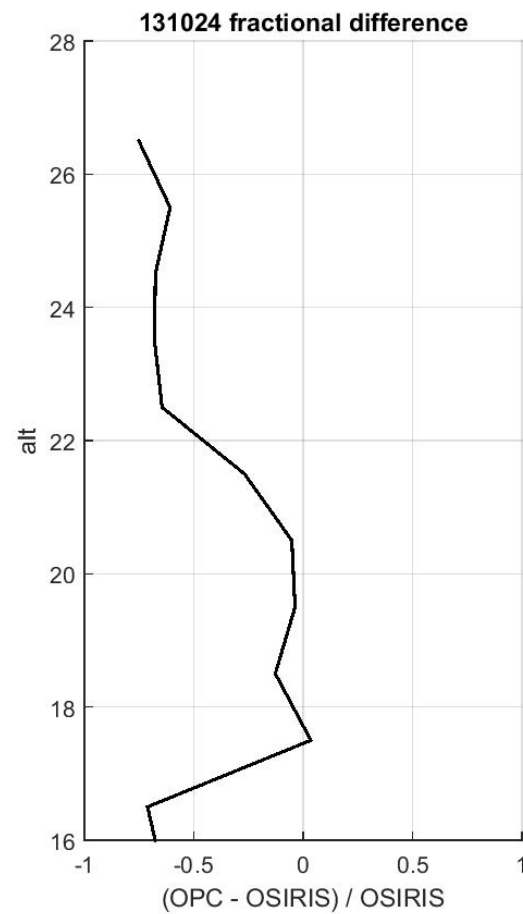
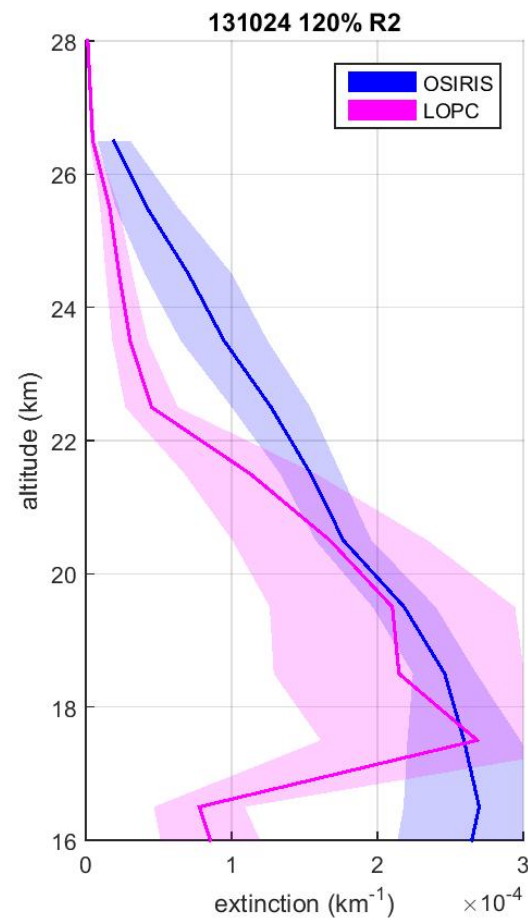
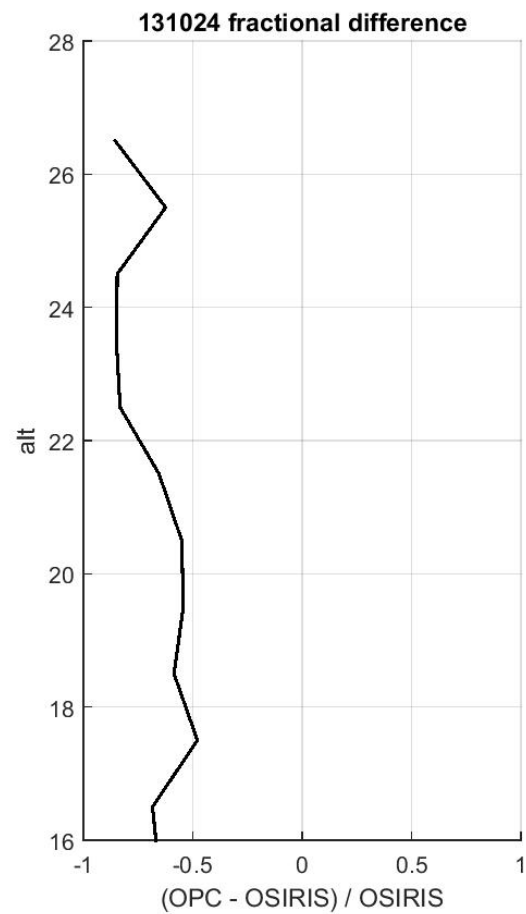
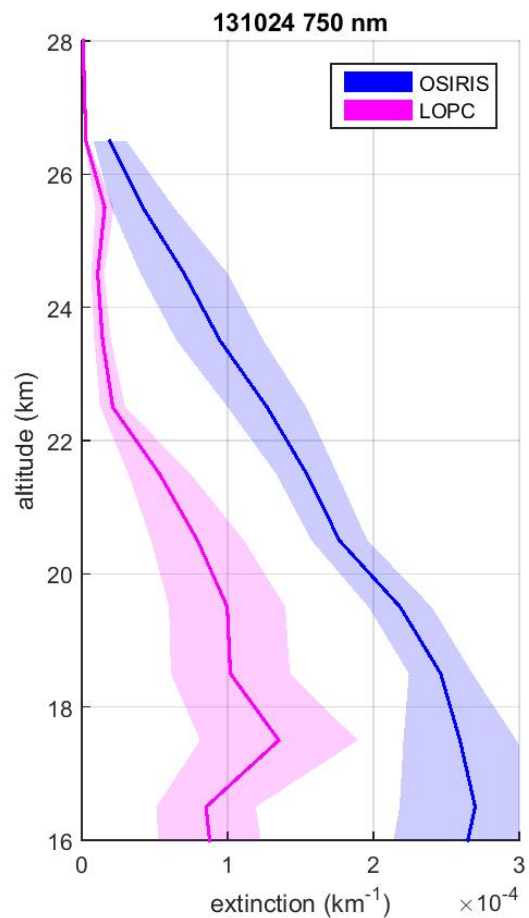
# OPC vs. OSIRIS





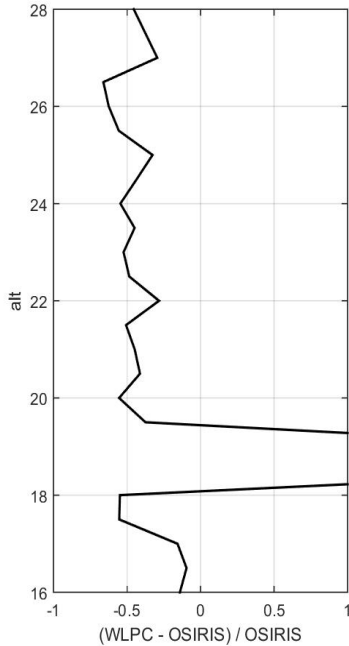
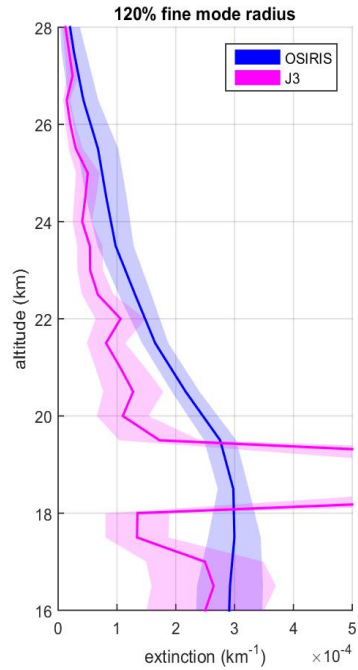
# OPC vs. OSIRIS

Increase coarse mode radius by 20%

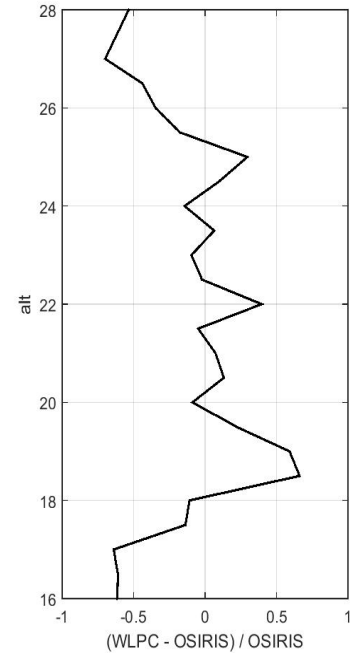
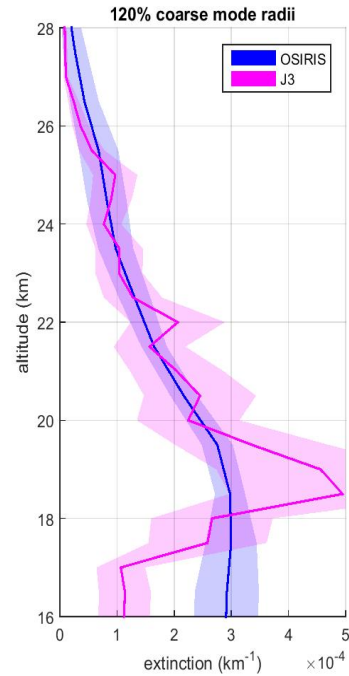


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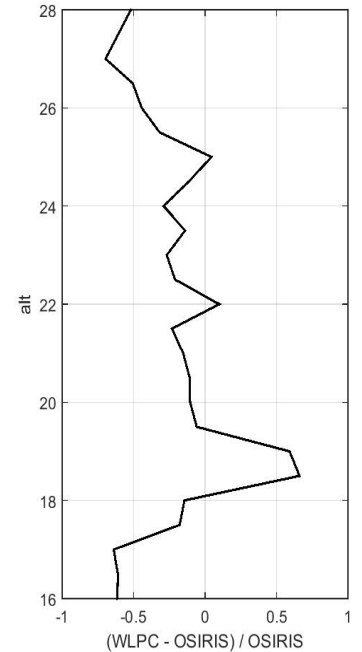
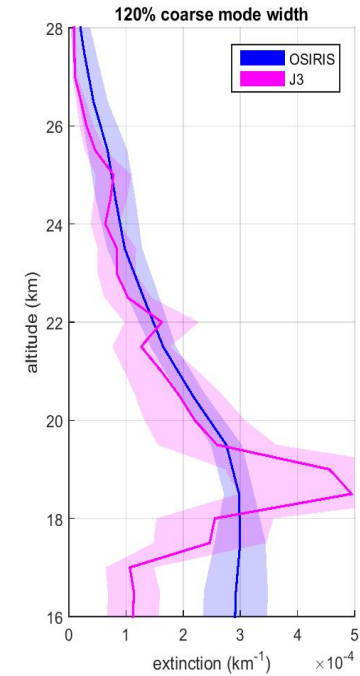
**Increase fine mode width by 20%**



**Increase coarse mode radius by 20%**



**Increase coarse mode width by 20%**



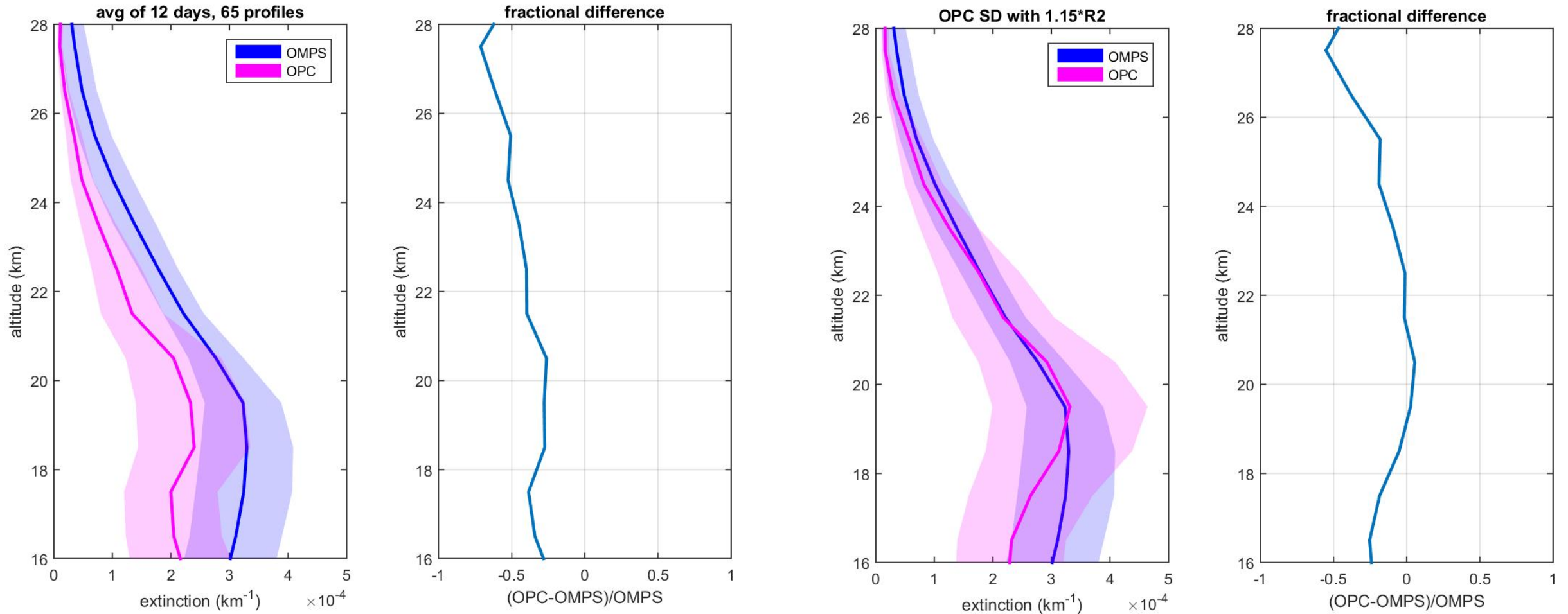
**Coarse mode more important than fine mode**

**Coarse mode radius =  $\sim 200$  nm**

**Fine mode radius =  $\sim 65$  nm**

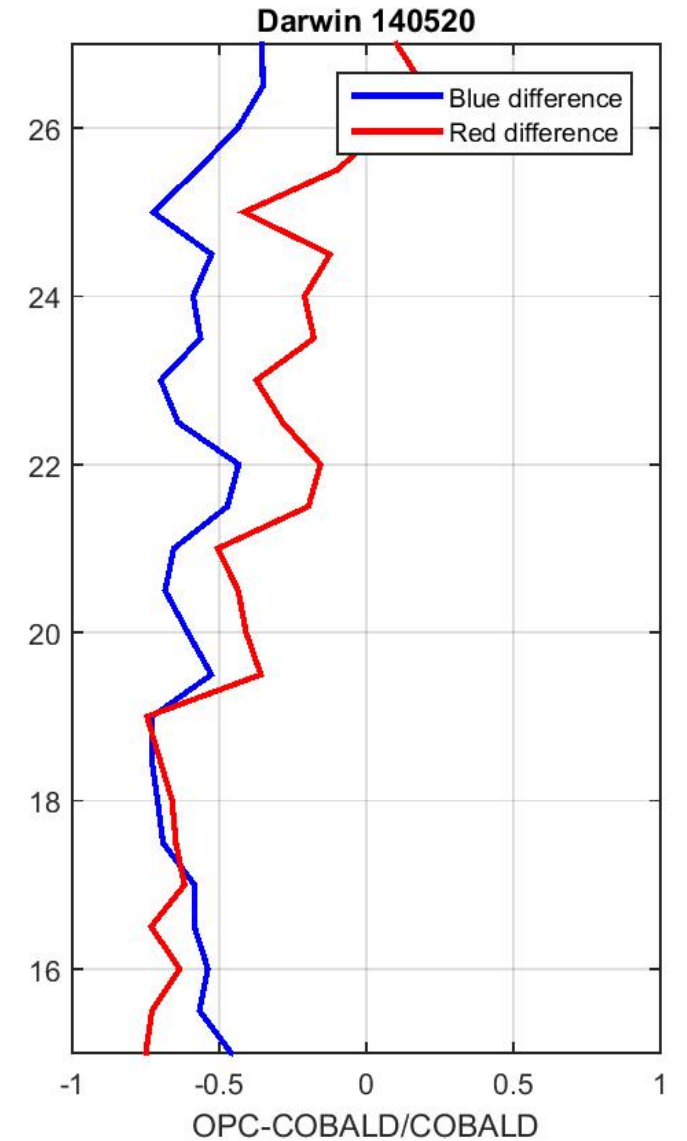
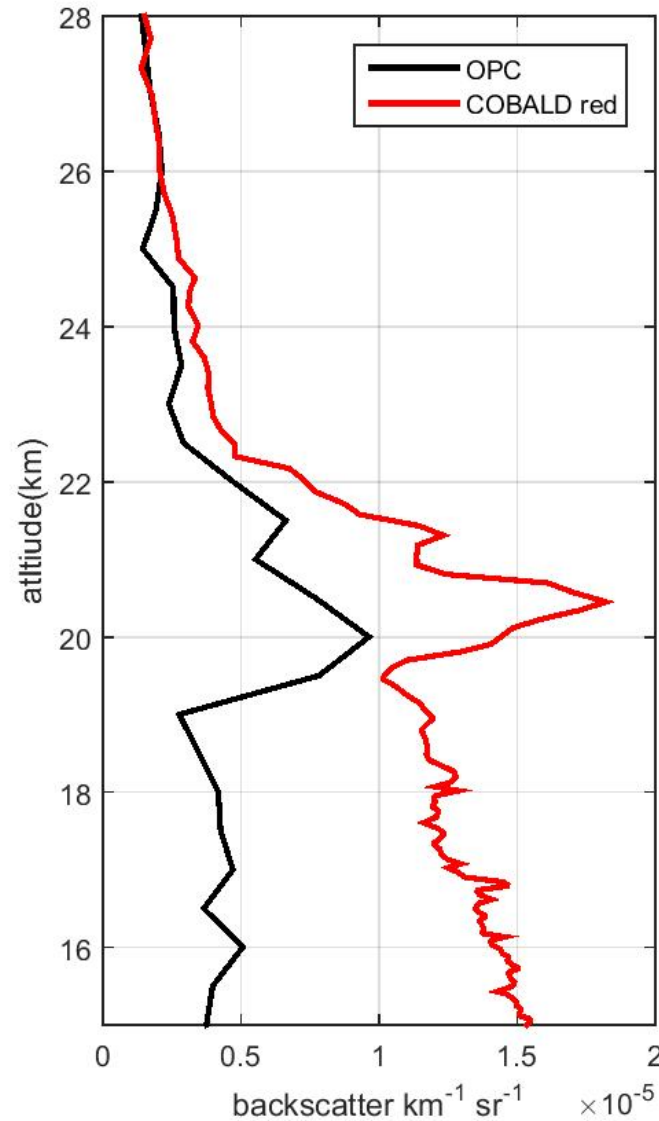
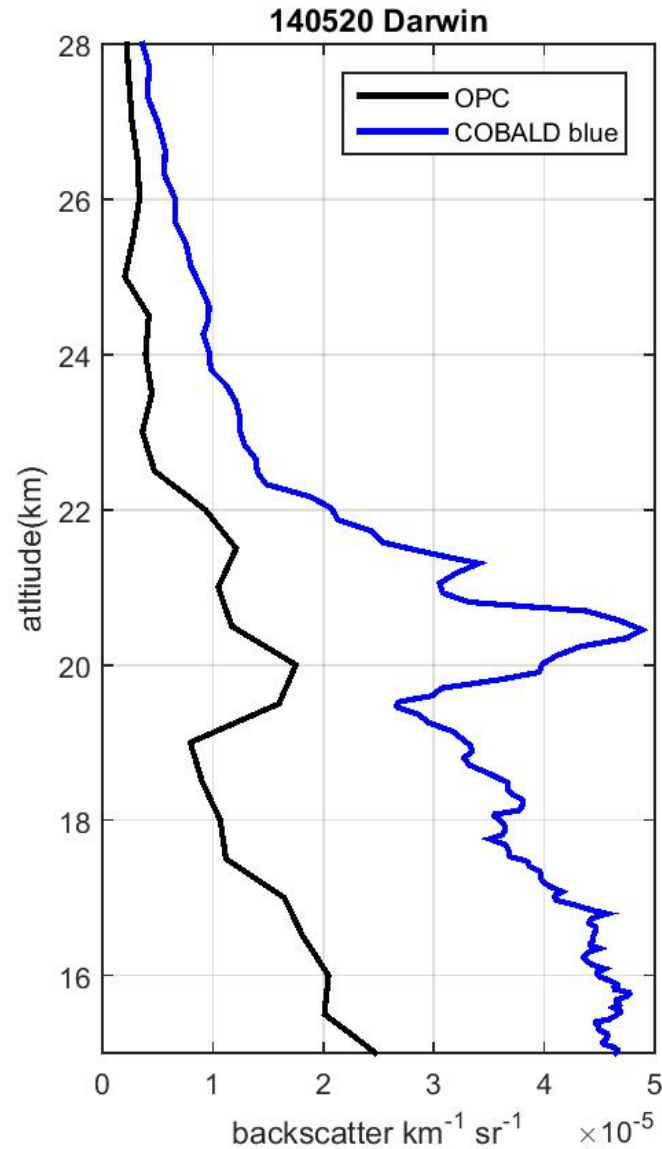
# Comparisons to OMPS

Increase coarse mode width by 15%

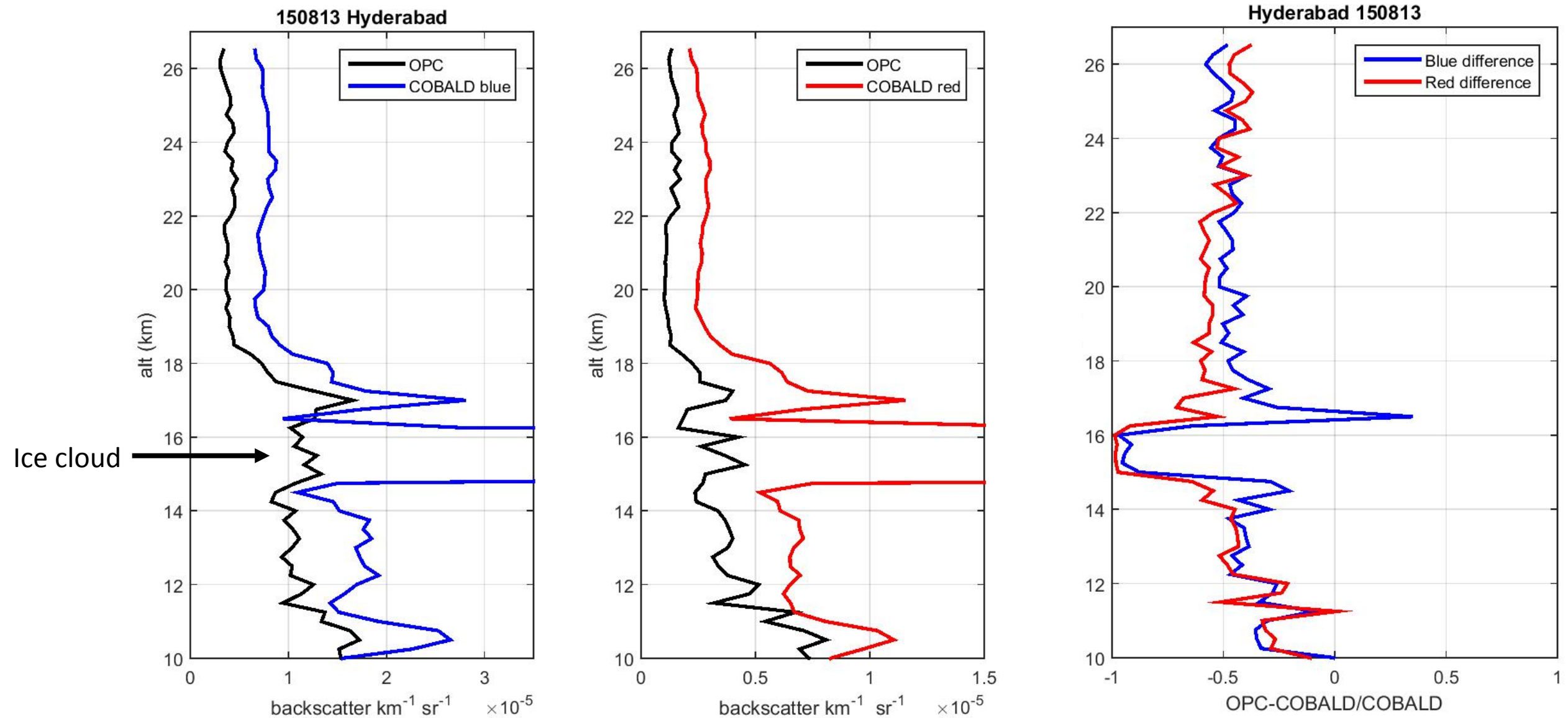


Coincident profiles from March 2012 - January 2015

# Comparisons to COBALD Backscatter-Sonde



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# Conclusions

No discernable  
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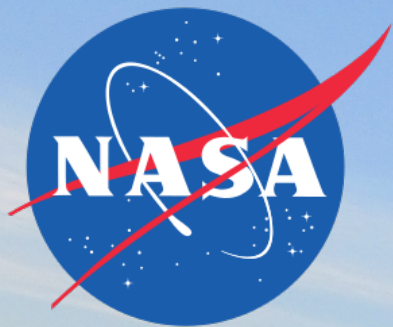
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2<sup>nd</sup> mode  
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Agreement with  
COBALD follows  
similar pattern





# Thank You



- OSIRIS team: Landon Rieger, Adam Bourassa
- COBALD team: Jean Paul Vernier, Frank Weinhold
- OMPS contact Ghassan Taha
- Terry Deshler

# Comparisons between 2 different UW OPCs

