

Koninklijk Belgisch Instituut voor Ruimte-Aeronomie (BIRA)

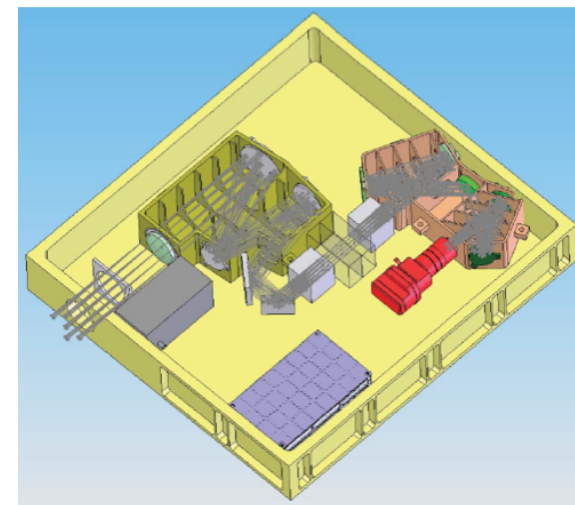
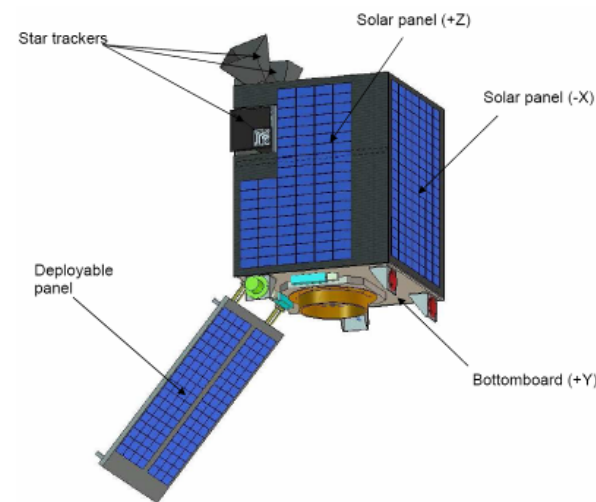
Overview of this presentation

The ALTIUS mission:

- What is it?
- Recent project developments
- Measurement modes
- Performance: simulated retrievals
- Measurement campaign results

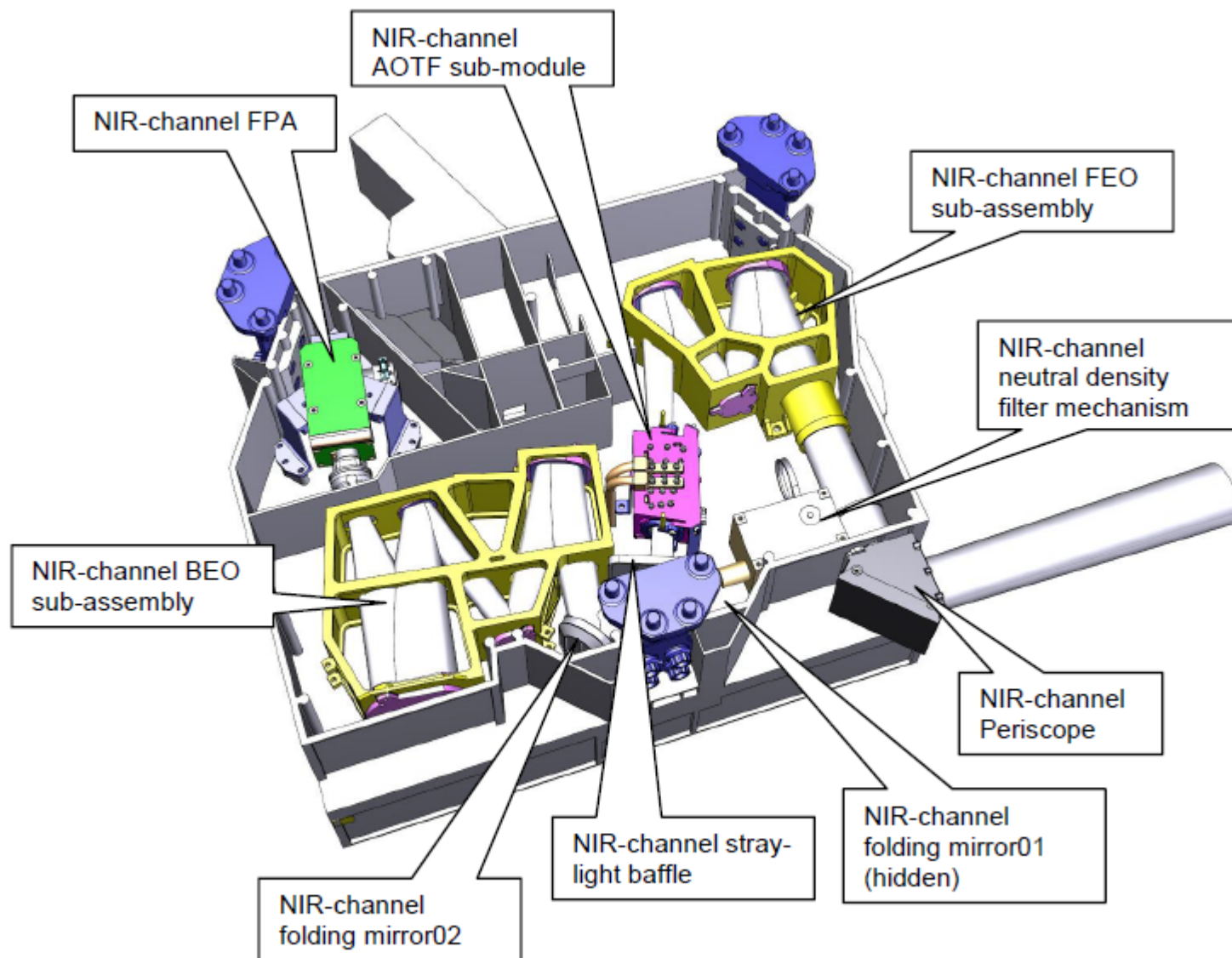
Atmospheric Limb Tracker for the Investigation of the Upcoming Stratosphere (ALTIUS)

- ALTIUS is a proposed **limb imager** (still in development phase)
- The instrument consists of three independent **spectral camera's** for the UV-Vis-NIR wavelength range (250–1800 nm).
- One individual spectral camera consists of optics + **AOTF** (Acousto-Optic Tunable Filter) + **imaging detector**. (for UV: Fabry-Perot filter)
- On board a **micro-satellite** (The Belgian **PROBA platform**) in **helio-synchronous** low Earth orbit
- Will operate in multi-mode — **limb scatter, solar/lunar/planetary/stellar occultation**.
- Main geophysical targets: strato-/mesospheric **ozone profiles and minor trace constituents** (NO_2 , H_2O , BrO , CH_4 , aerosols/clouds/PSCs/PMCs) and temperature (from refractive angle measurements)



ALTIUS instrument model at present: NIR channel

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ALTIUS: recent developments

(The ALTIUS-saga started already more than 10 years ago.)

- Industrial phase B1 completed (IDR) for the payload, platform completing its PDR.
- March 2015: ESA/EO proposed to make ALTIUS an ozone monitoring mission, BELSPO (Belgian Science Policy Office) funds the **OPEROZ study** (definition of requirements from scientific community and operational services)
- November 2015: ESA/EO review to check the compliance of ALTIUS to the OPEROZ study.
- December 2015: A scientific added value review by an independent panel of international experts was organized. (other atmospheric species)

Feb. 11 2016: Victory! Based on the two reviews, BELSPO expressed their intention to support the ALTIUS mission, welcoming participation by other EU Member States.

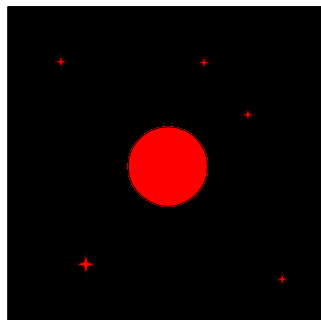
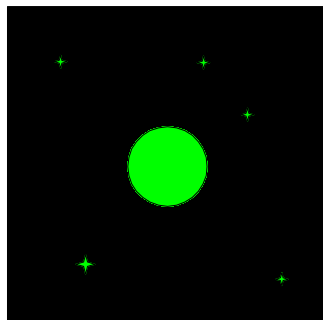
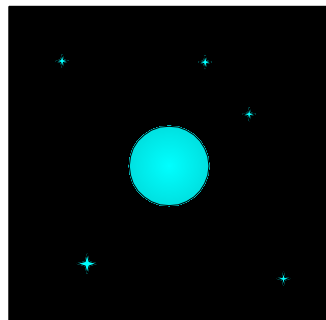
ALTIUS will most likely be implemented as an element in ESA (Earth Observation) Earth Watch Programme (monitoring missions).

Solar occultation: ALTIUS images

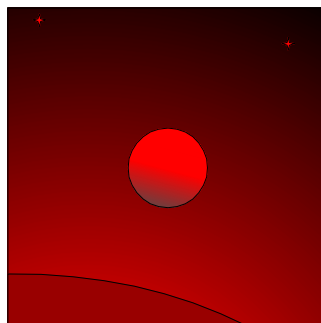
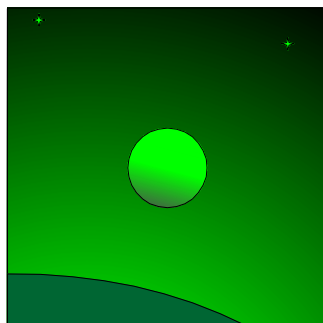
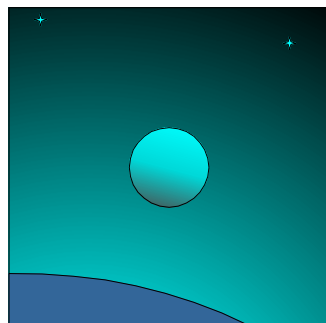
UV

Vis

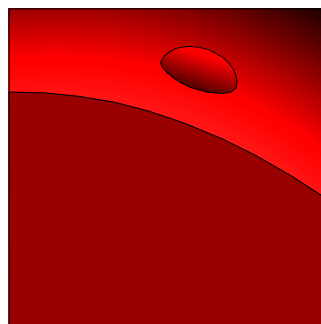
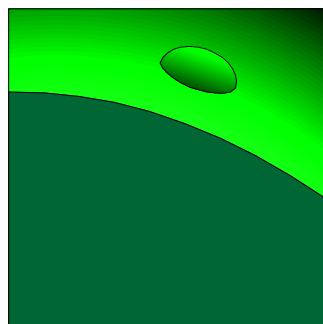
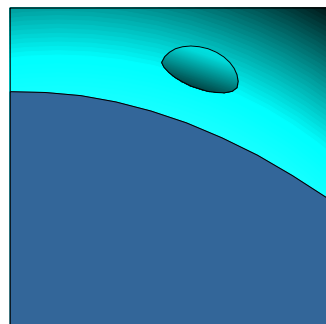
NIR



Outside
atmosphere



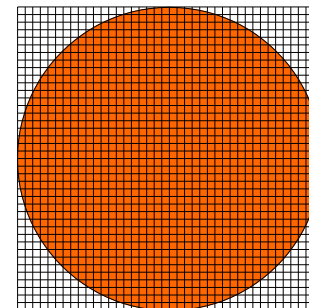
High
tangent
altitude



Low
tangent
altitude

- 2 occ/orbit
- 50 obs/occ of 1 sec
- 0.1 sec per image
- # images/obs: 4 per channel x 3 channels = 12
- Total time vertical occ: 50 sec
- Total # images per orbit: 1200 (cropping!)

1 pixel	0.2 mrad	0.6 km
Solar disk	9 mrad	27 km
Full solar disk = 45 x 45 pixels		

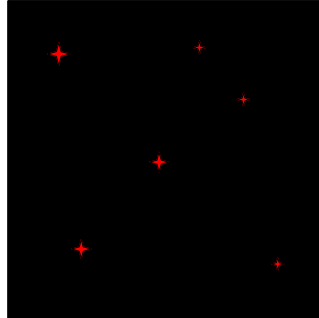
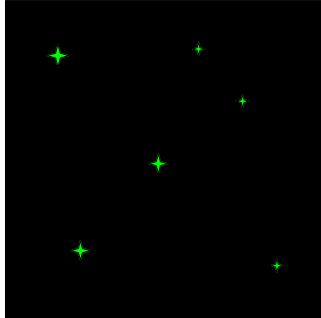
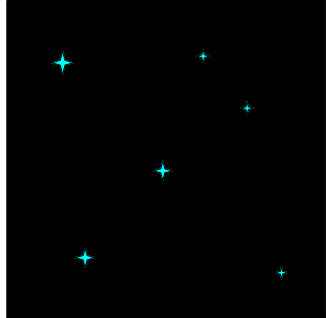


Stellar occultation: ALTIUS images

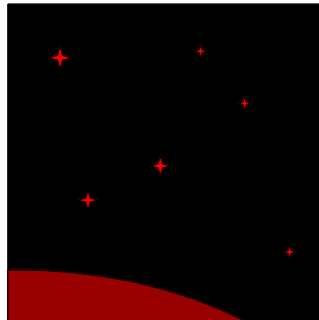
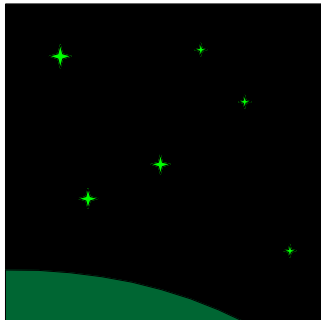
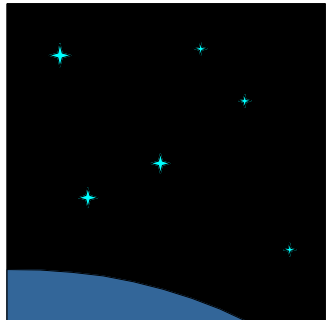
UV

Vis

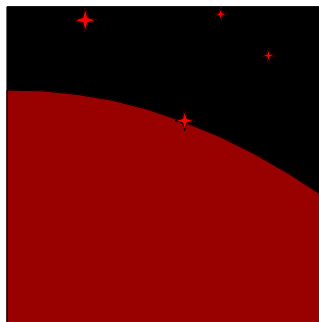
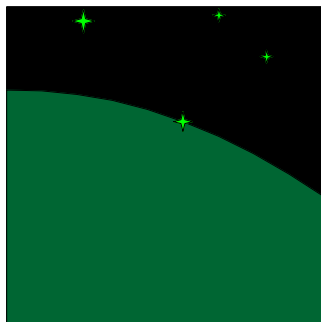
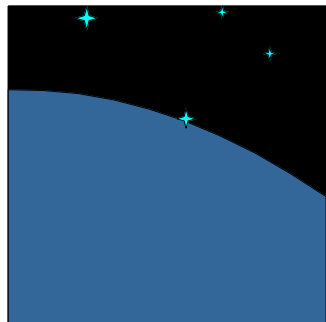
NIR



Outside
atmosphere



High
tangent
altitude

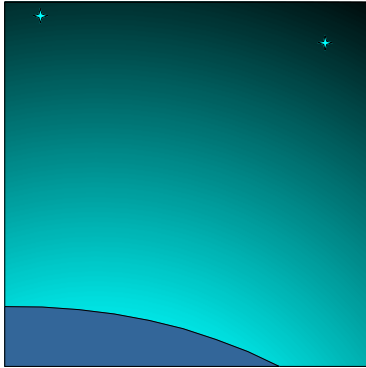


Low
tangent
altitude

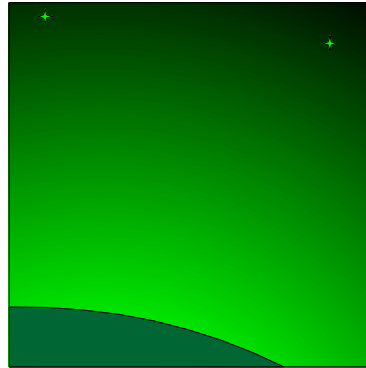
- Occultation of several stars possible (but rare).
- typically about 5 occultations per orbit
- 50 obs per occ of 2 sec
- 0.5 sec per image
- # images per observation: 4 per channel x 3 channels = 12
- Total # images per orbit: 6000

Limb observation: ALTIUS images

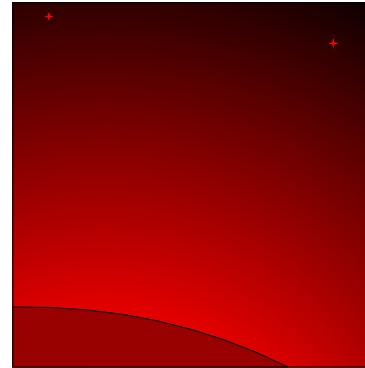
UV



Vis



NIR



All
altitudes
at once!

- Simplest observation mode, very complex forward model!
- About 50 observations per orbit
- # images per observation: 4 per channel x 3 channels = 12
- 1 second per image
- About 10 seconds per observation
- Total # images per orbit: about 600

A typical measurement sequence (1 orbit)

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Event description	Time (s)	Sat Lat	Sat Lon	TP Lat	TP Lon	Event description	Time (s)	Sat Lat	Sat Lon	TP Lat	TP Lon	Event description	Time (s)	Sat Lat	Sat Lon	TP Lat	TP Lon
SoO (sunrise)	0	64.7	-65.1	70.8	-117.1	BL Obs	1916	-3.0	124.0	20.5	129.1	BL Obs	3056	-70.8	95.5	-48.2	112.1
	66	68.4	-69.0	71.0	-137.8	rest	1926	-3.6	123.8	19.9	128.9	BL Obs	3066	-71.4	94.7	-48.8	111.9
Minoraune						BL Obs	1976	-6.7	123.2	16.9	128.2	BL Obs	3116	-74.1	89.5	-51.8	110.6
BL Obs	836	61.8	144.3	81.3	-161.5	rest	1986	-7.3	123.1	16.3	128.1	BL Obs	3126	-74.6	88.3	-52.4	110.4
	846	61.3	143.8	81.0	-165.2	BL Obs	2036	-10.3	122.4	13.2	127.4	rest	3176	-77.1	80.9	-55.3	108.9
rest						rest	2046	-10.9	122.3	12.6	127.3	BL Obs	3186	-77.6	79.1	-55.9	108.6
BL Obs	896	58.3	141.9	79.4	-8.4	rest	2096	-13.9	121.6	9.6	126.6	rest	3236	-79.7	67.7	-58.9	106.9
	906	57.7	141.6	79.0	176.9	BL Obs	2106	-14.5	121.5	9.0	126.5	BL Obs	3246	-80.0	64.9	-59.5	106.6
rest						rest	2156	-17.5	120.8	6.0	125.9	rest	3296	-81.5	47.8	-62.4	104.5
BL Obs	956	54.8	140.0	76.7	166.9	BL Obs	2166	-18.1	120.7	5.4	125.7	BL Obs	3306	-81.6	43.8	-63.0	104.1
	966	54.2	139.7	76.2	165.3	rest	2216	-21.1	120.0	2.4	125.1	rest	3356	-81.9	22.3	-65.8	101.6
rest						BL Obs	2226	-21.7	119.9	1.8	125.0	BL Obs	3366	-81.8	18.0	-66.4	101.0
BL Obs	1016	51.2	138.3	73.6	158.7	rest	2276	-24.8	119.1	-1.3	124.3	rest	3416	-80.8	-1.2	-69.2	97.8
	1026	50.6	138.1	73.1	157.6	BL Obs	2286	-25.4	119.0	-1.9	124.2	BL Obs	3426	-80.4	-4.5	-69.8	97.0
rest						rest	2336	-28.4	118.2	-4.9	123.6	rest	3476	-78.5	-17.7	-72.5	92.7
BL Obs	1076	47.6	136.8	70.4	153.0	BL Obs	2346	-29.0	118.1	-5.5	123.4	BL Obs	3486	-78.1	-19.9	-73.1	91.7
	1086	47.0	136.6	69.8	152.2	rest	2396	-32.0	117.3	-8.5	122.8	rest	3536	-75.7	-28.5	-75.7	85.5
rest						rest	2406	-32.6	117.1	-9.1	122.7	BL Obs	3546	-75.2	-29.9	-76.2	84.0
BL Obs	1136	44.0	135.5	67.0	148.8	BL Obs	2456	-35.6	116.3	-12.1	122.0	rest	3596	-72.6	-35.7	-78.5	74.8
	1146	43.4	135.3	66.4	148.2	rest	2466	-36.2	116.1	-12.7	121.9	BL Obs	3606	-72.0	-36.7	-78.9	72.5
rest						BL Obs	2516	-39.2	115.2	-15.8	121.2	rest	3656	-69.3	-40.8	-80.7	58.4
BL Obs	1196	40.4	134.3	63.5	145.6	rest	2526	-39.8	115.1	-16.4	121.1	BL Obs	3666	-68.7	-41.5	-81.0	54.9
	1206	39.8	134.1	63.0	145.1	BL Obs	2576	-42.8	114.1	-19.4	120.4	rest	3716	-65.9	-44.6	-81.9	35.0
rest						rest	2586	-43.4	113.9	-20.0	120.3	BL Obs	3726	-65.3	-45.2	-81.9	30.6
BL Obs	1256	36.8	133.2	60.0	143.0	rest	2636	-46.4	112.8	-23.0	119.6	rest	3833	-57.9	-50.6	-78.5	-78.8
	1266	36.2	133.1	59.4	142.6	BL Obs	2646	-47.0	112.6	-23.6	119.4	BL Obs	3918	-54.1	-52.7	-73.6	-74.3
rest						rest	2696	-49.9	111.4	-26.6	118.7	Minoraune					
BL Obs	1316	33.2	132.2	56.5	140.9	BL Obs	2706	-50.5	111.2	-27.2	118.5	StQ M-H0.4, Hyp-37279	4188	-38.0	-58.7	-38.9	-30.1
	1326	32.6	132.1	55.9	140.6	rest	2756	-53.5	109.8	-30.2	117.8	Minoraune	4396	-26.0	-61.9	-26.9	-40.8
rest						BL Obs	2766	-54.1	109.5	-30.8	117.6	Minoraune	4656	-9.7	-65.6	-31.7	-68.8
BL Obs	1376	29.6	131.3	52.9	139.1	rest	2816	-57.1	108.0	-33.8	116.8	StQ M-L59 Hyp-61084	4722	-5.7	-66.5	-24.5	-68.3
	1386	29.0	131.1	52.3	138.8	BL Obs	2826	-57.6	107.7	-34.4	116.6	Minoraune	4752	-3.9	-66.9	-26.0	-68.3
rest						rest	2876	-60.6	105.8	-37.4	115.8	StQ M-H0.77, Hyp-60718	4818	0.1	-67.7	-18.7	-69.9
BL Obs	1436	26.0	130.4	49.4	137.5	rest	2886	-61.2	105.4	-38.0	115.6	Minoraune	4866	3.0	-68.3	-17.6	-77.0
	1446	25.4	130.2	48.8	137.3	BL Obs	2936	-64.1	103.2	-41.0	114.7	StQ M-L44, Hyp-32349	4888	7.3	-68.3	-9.9	-76.7
rest						rest	2946	-64.6	102.7	-41.6	114.5	Minoraune	5053	14.3	-70.8	8.6	-48.9
BL Obs	1486	22.4	129.5	45.8	136.1	BL Obs	2986	-67.5	99.9	-44.6	113.5	rest	5262	26.9	-73.7	23.4	-53.3
	1506	21.7	129.3	45.2	135.9	rest	3006	-68.1	99.2	-45.2	113.2	Minoraune	5365	33.1	-75.3	39.6	-53.0
rest						BL Obs						StQ M-H0.4, Hyp-37279	5368	45.3	-79.1	53.6	-47.2
BL Obs	1556	18.7	128.6	42.2	134.9	rest						Minoraune	5655	51.1	-81.4	70.4	-68.1
	1566	18.1	128.5	41.6	134.7	rest						StQ M-L36 Hyp-49569	5796	55.4	-83.4	76.5	-66.0
rest																	
BL Obs	1616	15.1	127.8	38.6	133.8												
	1626	14.5	127.7	38.0	133.6												
rest																	
BL Obs	1676	11.5	127.0	35.0	132.7												
	1686	10.9	126.9	34.4	132.5												
rest																	
BL Obs	1736	7.8	126.3	31.3	131.7												
	1746	7.2	126.1	30.7	131.6												
rest																	
BL Obs	1796	4.2	125.5	27.7	130.8												
	1806	3.6	125.4	27.1	130.6												
rest																	
BL Obs	1856	0.6	124.7	24.1	129.9												
	1866	0.0	124.6	23.5	129.8												
rest																	

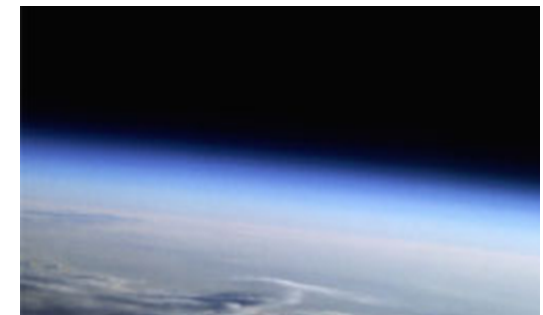
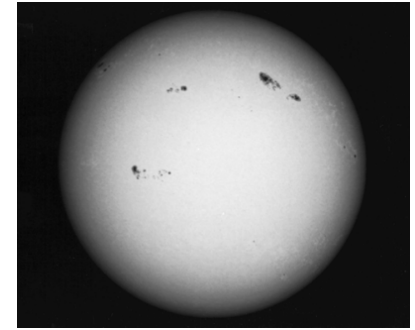
Non-nominal observation modes

- Multi-orbit tomography:
 - backward-forward + left-looking in next orbit
- Nadir:
 - 1 pixel = 140 x 140 meter
 - Total FOV: about 23 x 23 km (= scale of a city)
- Planet occultation:
 - same principle as stellar occultation
- Lunar occultation:
 - same principle as solar occultation
 - Albedo!
- Observations of optical emissions (specific wavelength tuning)
 - Aurora (555.7 nm green line)
 - nightglow (auroral green line, red lines)
- Calibration campaigns
- Special campaigns: volcanic eruptions



Simulations: forward models

- **Solar occultation:**
 - Beer-Lambert extinction law
 - Solar limb darkening
 - Refractive effects (disk flattening, dilution, refractive bending)
 - ALTIUS instrument functions (PSF, spectral response, integration time)
- **Stellar occultation:** same as solar, except:
 - Various stellar magnitudes (S/N varies)
 - Various star temperatures (spectral emission varies)
 - Point sources (No angular integration)
- **Limb observations:**
 - Full radiative transfer code (“ARTS”:ALTIUS Radiative Transfer Simulator)
 - Spherical geometry
 - Vectorial (Stokes vector as output)
 - ALTIUS instrument functions



Occultation simulations: parameters

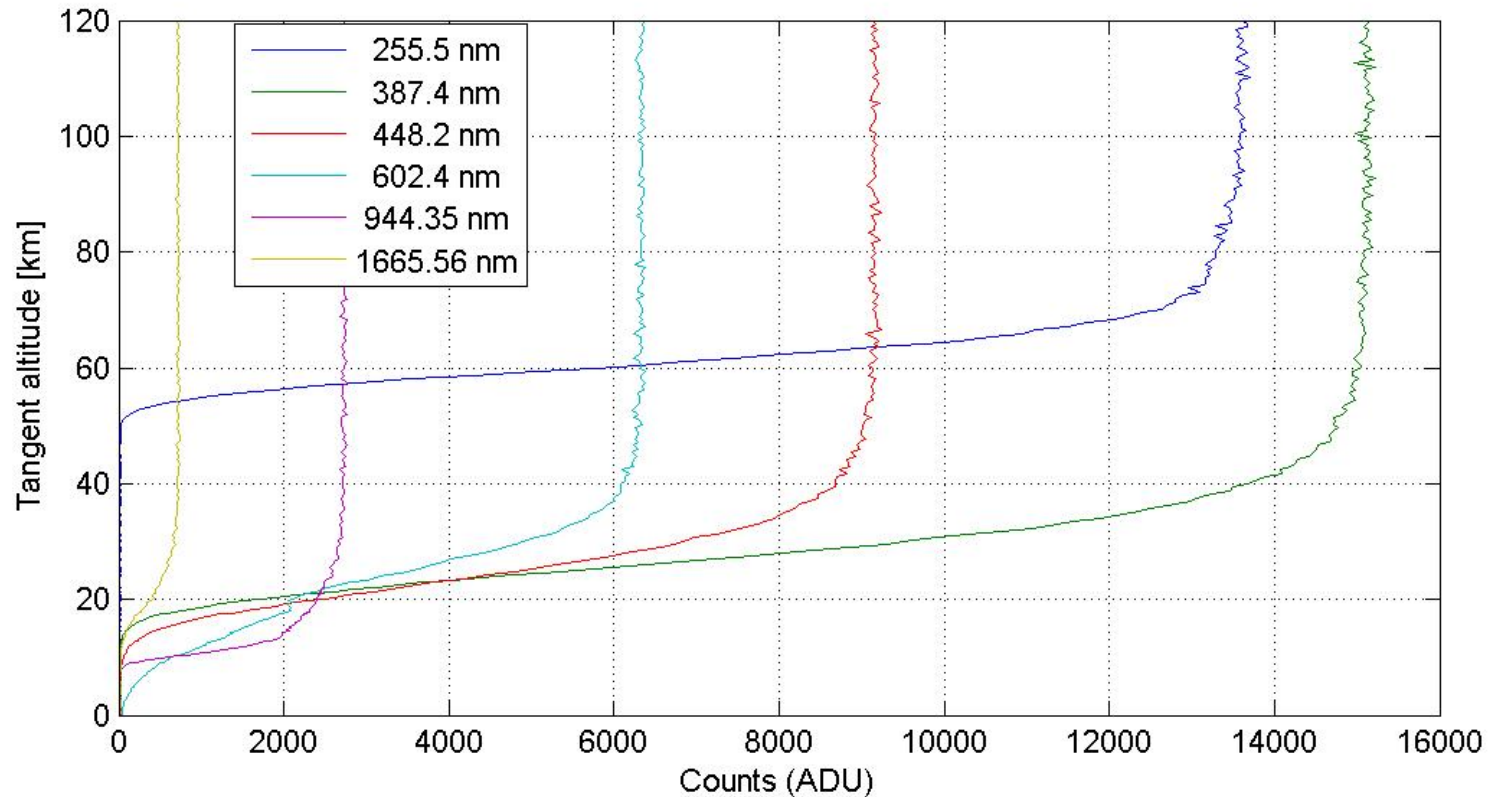
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	FOV	Nb pixels	Vertical sampling	Aperture diameter	Spectral range
UV (solar)	2°	170x170	600m	9 mm	250-450 nm
UV (stellar)	0.2°			50 mm	
VIS	2°	512x512	200m	41 mm	440-800 nm
NIR	2°	256x256	400m	41 mm	800-1800 nm

Geometrical parameters	
Day of the year	# 180
Tangent point latitude	50°N
LOS azimuth angle	180° (backward looking; sunset/starset)
Local Earth radius	6366 km (from the WGS84 model)
Satellite orbital altitude	650 km

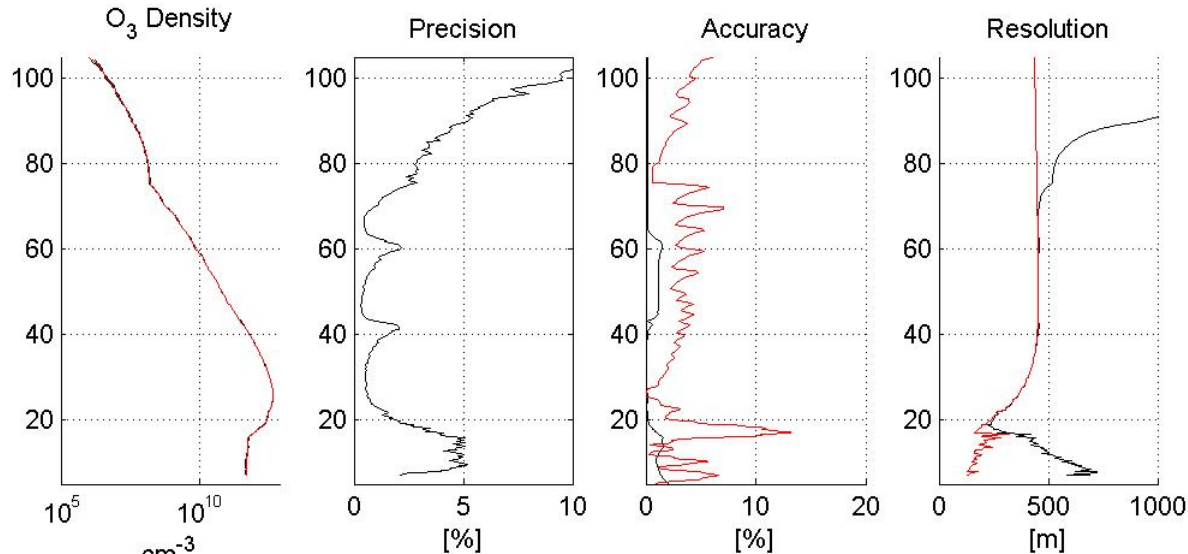
Channel	Wavelength [nm]	Constituent
UV	255.5	Ozone (Hartley band peak)
	300	Ozone (Hartley band slope)
	338.4	BrO absorption line peak
	387.4	OCIO absorption line peak
Visible	448.2	NO ₂ absorption line (peak)
	450.0	NO ₂ absorption line (valley)
	525.0	O ₃ Chappuis band (slope)
	602.4	O ₃ Chappuis band (peak)
Near-infrared	944.35	H ₂ O absorption band
	1020.0	Stratospheric and tropospheric aerosols
	1070.0	Stratospheric and tropospheric aerosols
	1665.56	CH ₄ absorption band (Q-branch)

Solar occultation: simulated signals

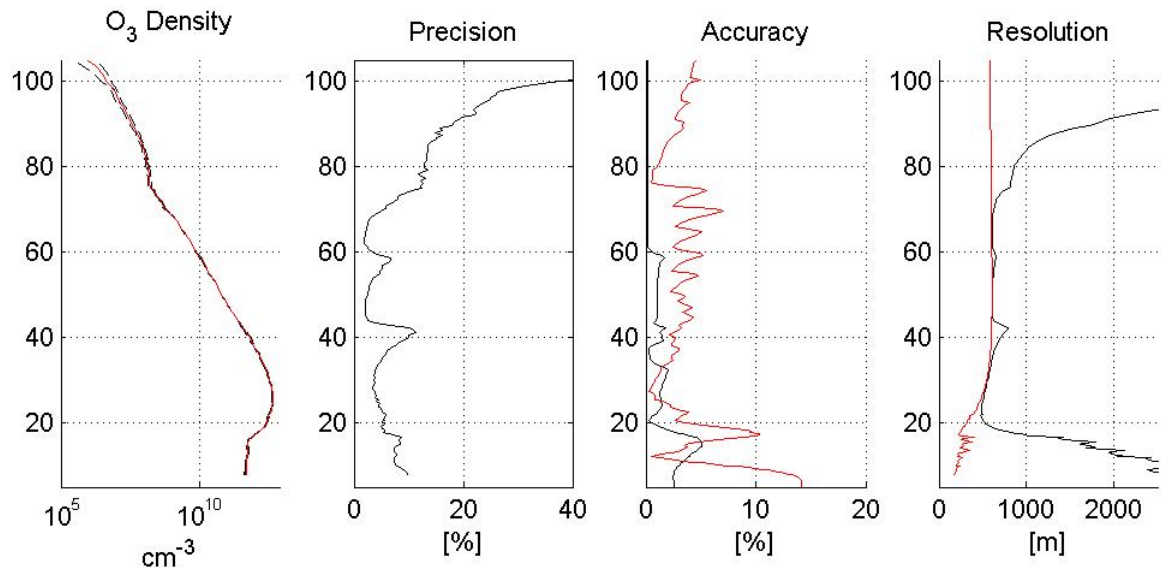


- Noise contributions: Poisson shot noise, digitization error (14 bit)
- Additional errors (according to instrument requirements):
 - Altitude pointing error: 150 m
 - Center of spectral bandwidth: $2\sigma = 0.1 \text{ nm}$

Occultation results: midlatitude ozone



Solar occultation



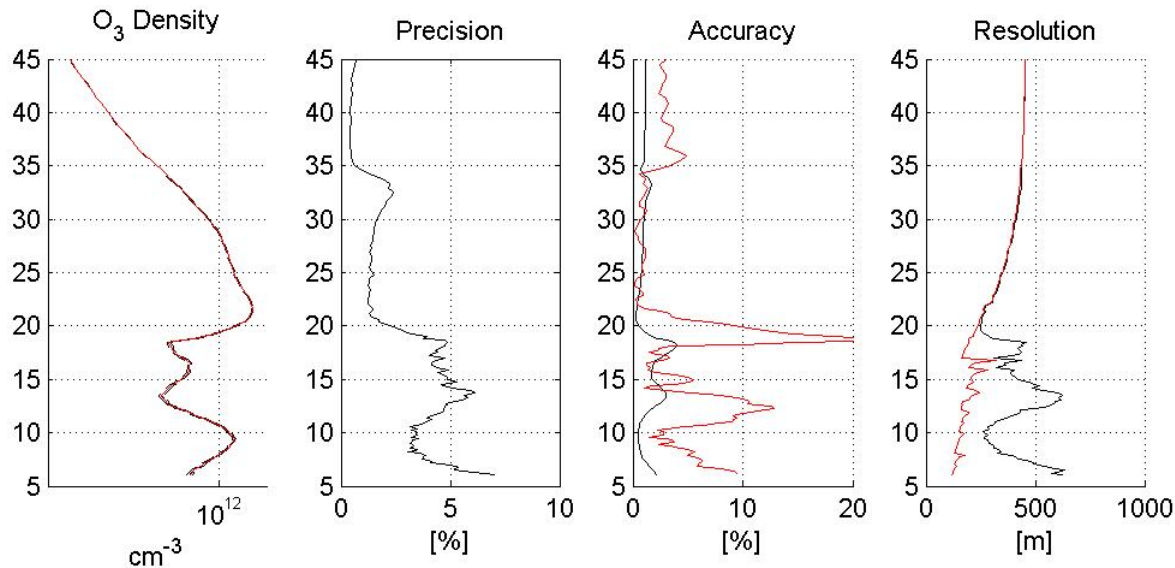
Stellar occultation:

- Sirius (α CMa)
- Magnitude: -1.44
- Temp: 11000 K

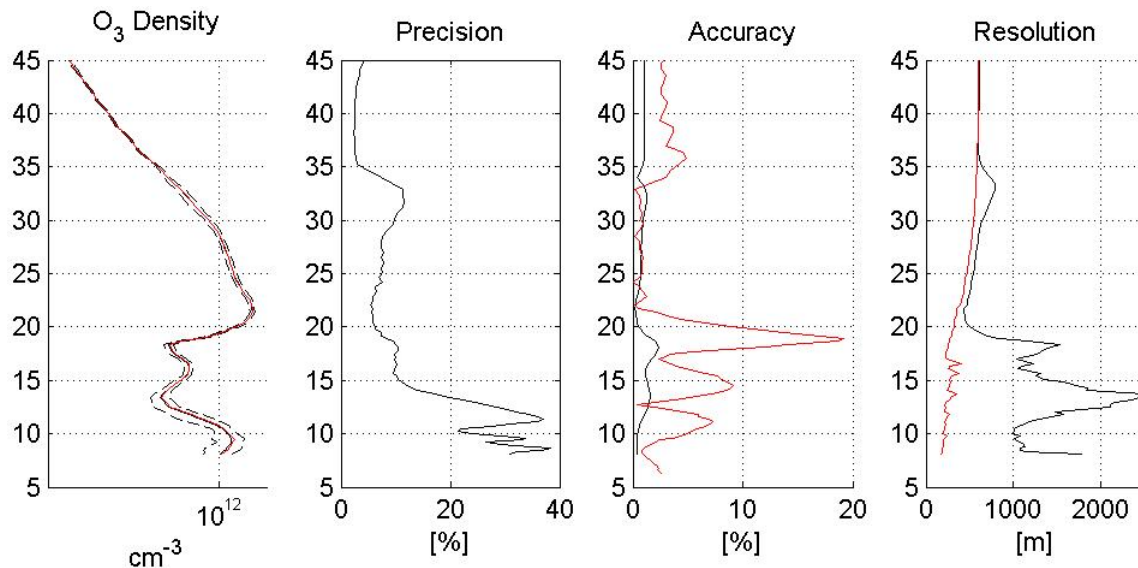
Important error due to wrong pointing

Occult. results: polar ozone, Antarctic spring

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Solar occultation



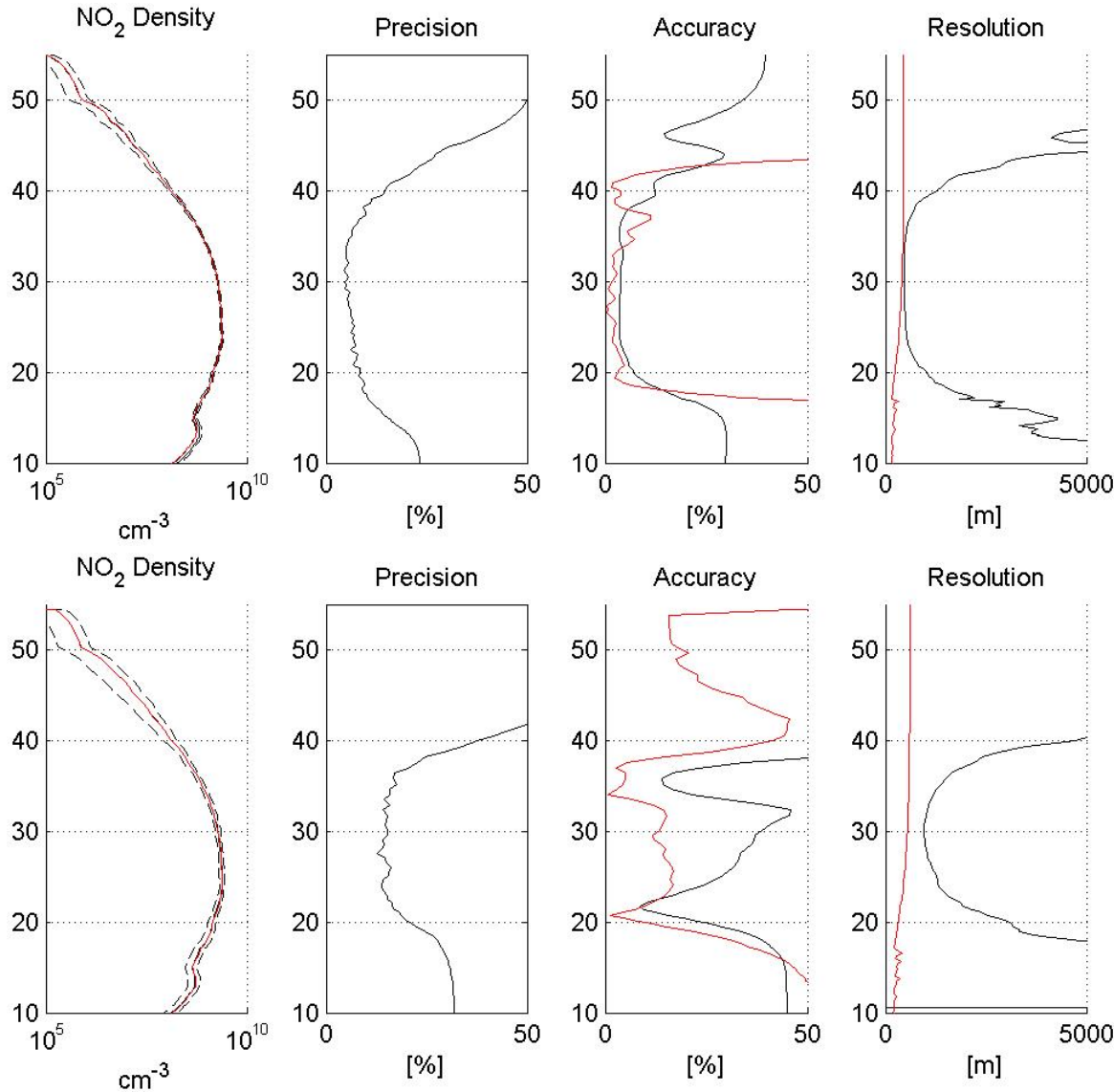
Stellar occultation:

- Sirius (αCMa)
- Magnitude: -1.44
- Temp: 11000 K

Increased error due to pointing at the top and bottom of the ozone hole

Occultation results: NO₂

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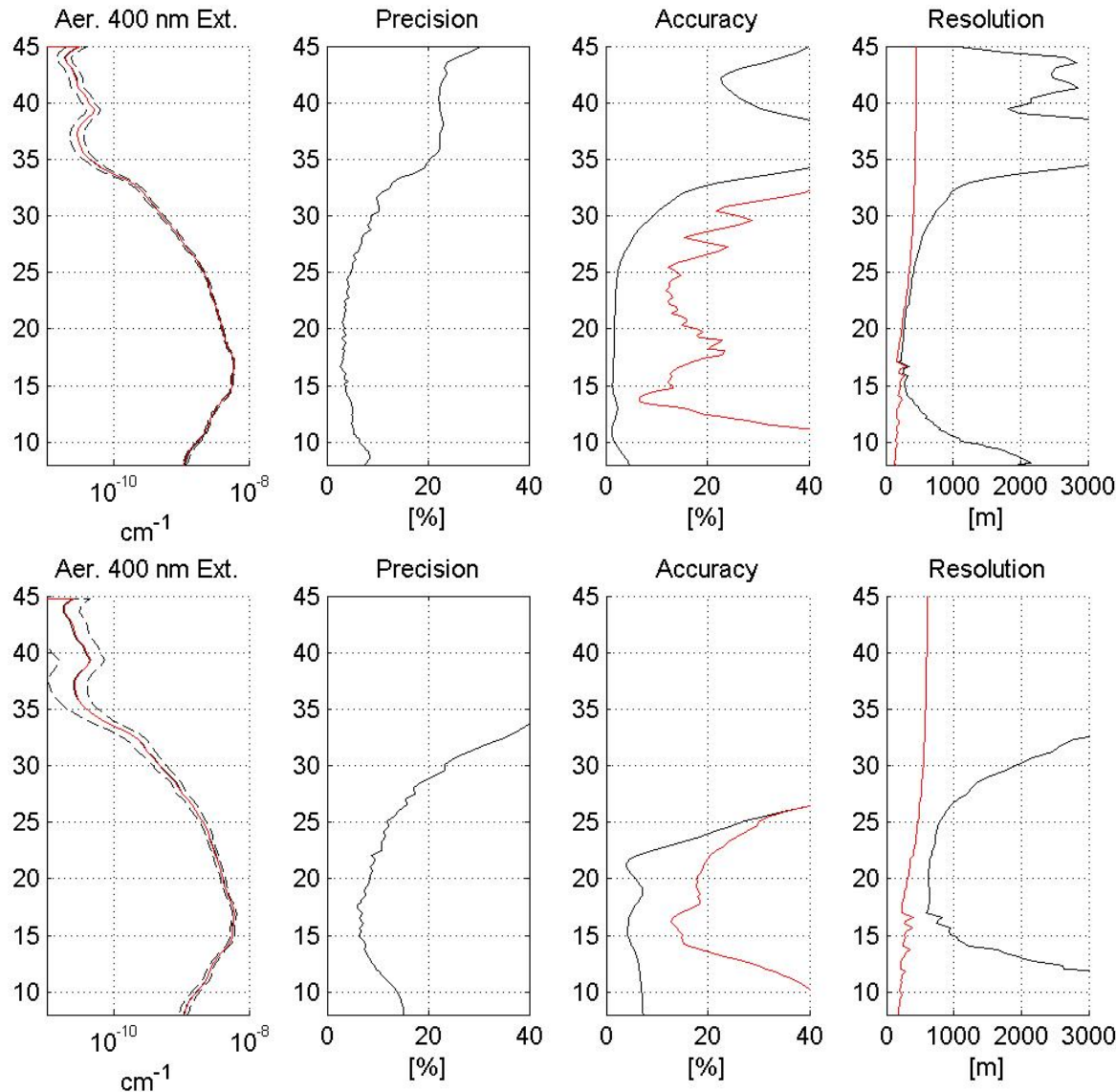
Solar occultation

Stellar occultation:

- Sirius (αCMa)
- Magnitude: -1.44
- Temp: 11000 K

*For stellar occ.,
wavelength assignment
error is very pronounced*

Occultation results: Aerosol ext. at 400 nm



Solar occultation

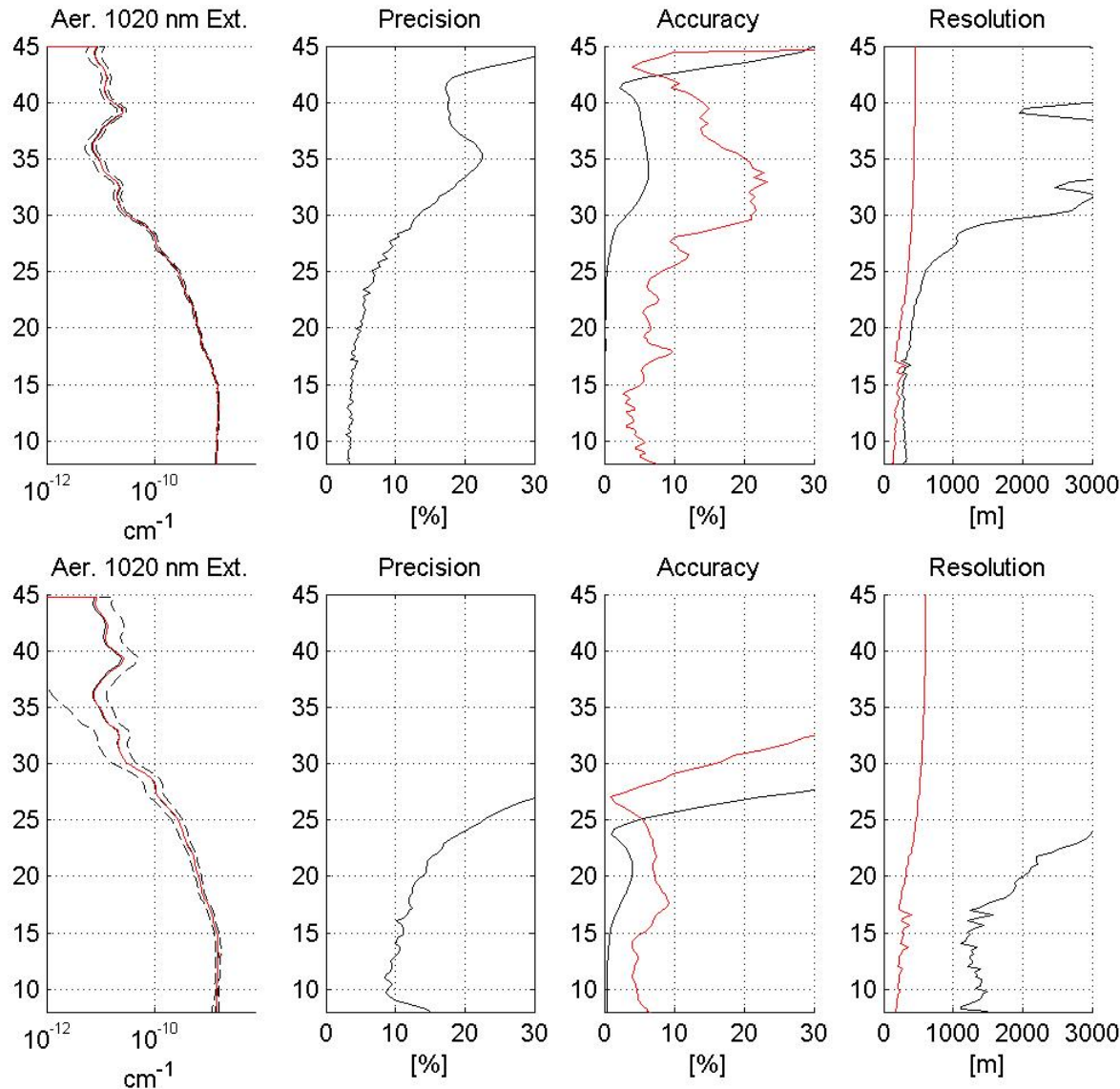
Stellar occultation:

- Sirius (α CMa)
- Magnitude: -1.44
- Temp: 11000 K

*Main error source:
pointing error*

Occultation results: Aerosol ext. at 1020 nm

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Solar occultation

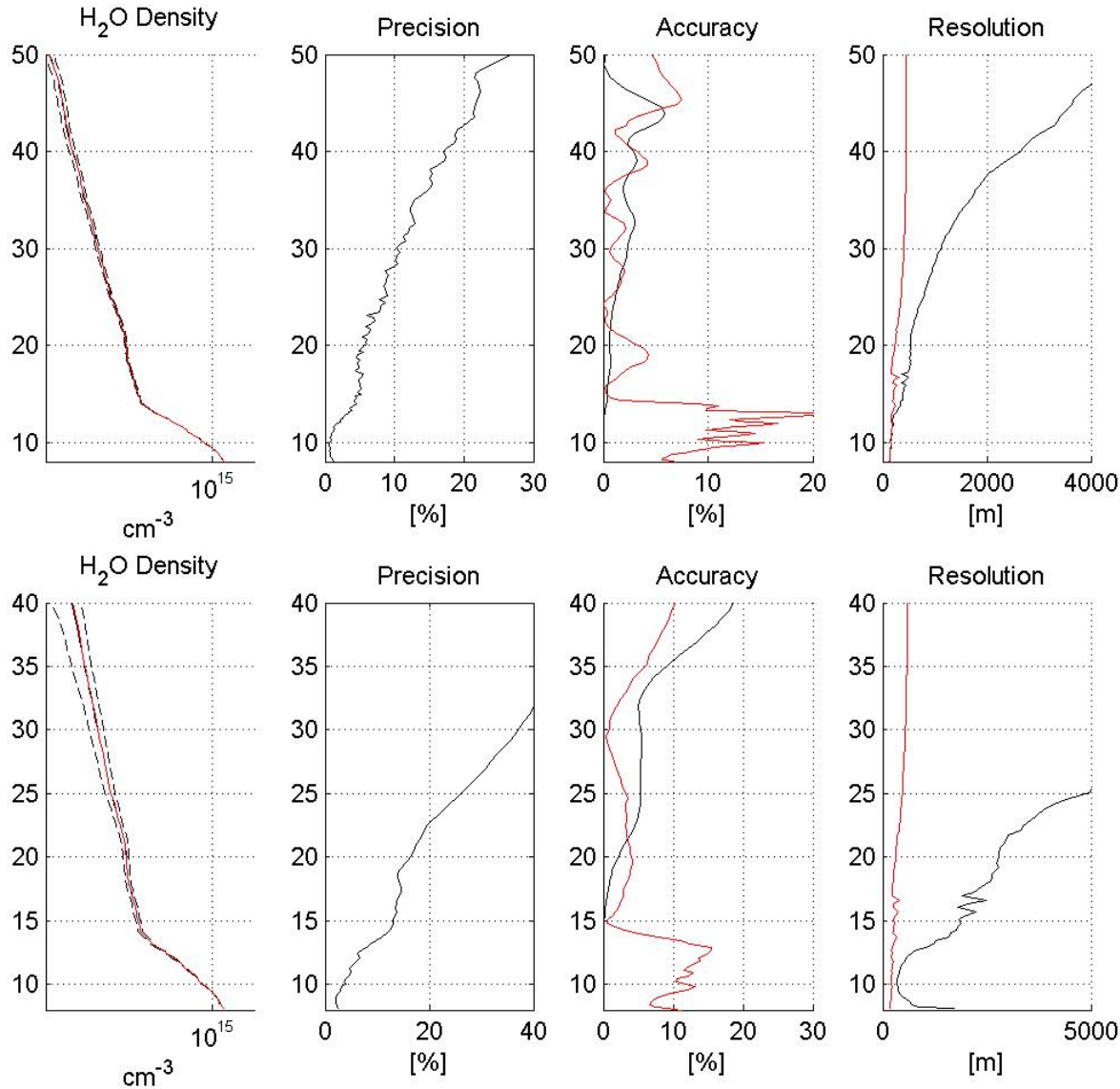
Stellar occultation:

- Sirius (α CMa)
- Magnitude: -1.44
- Temp: 11000 K

*Main error sources:
pointing error and
detector noise*

Occultation results: water vapour

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Solar occultation

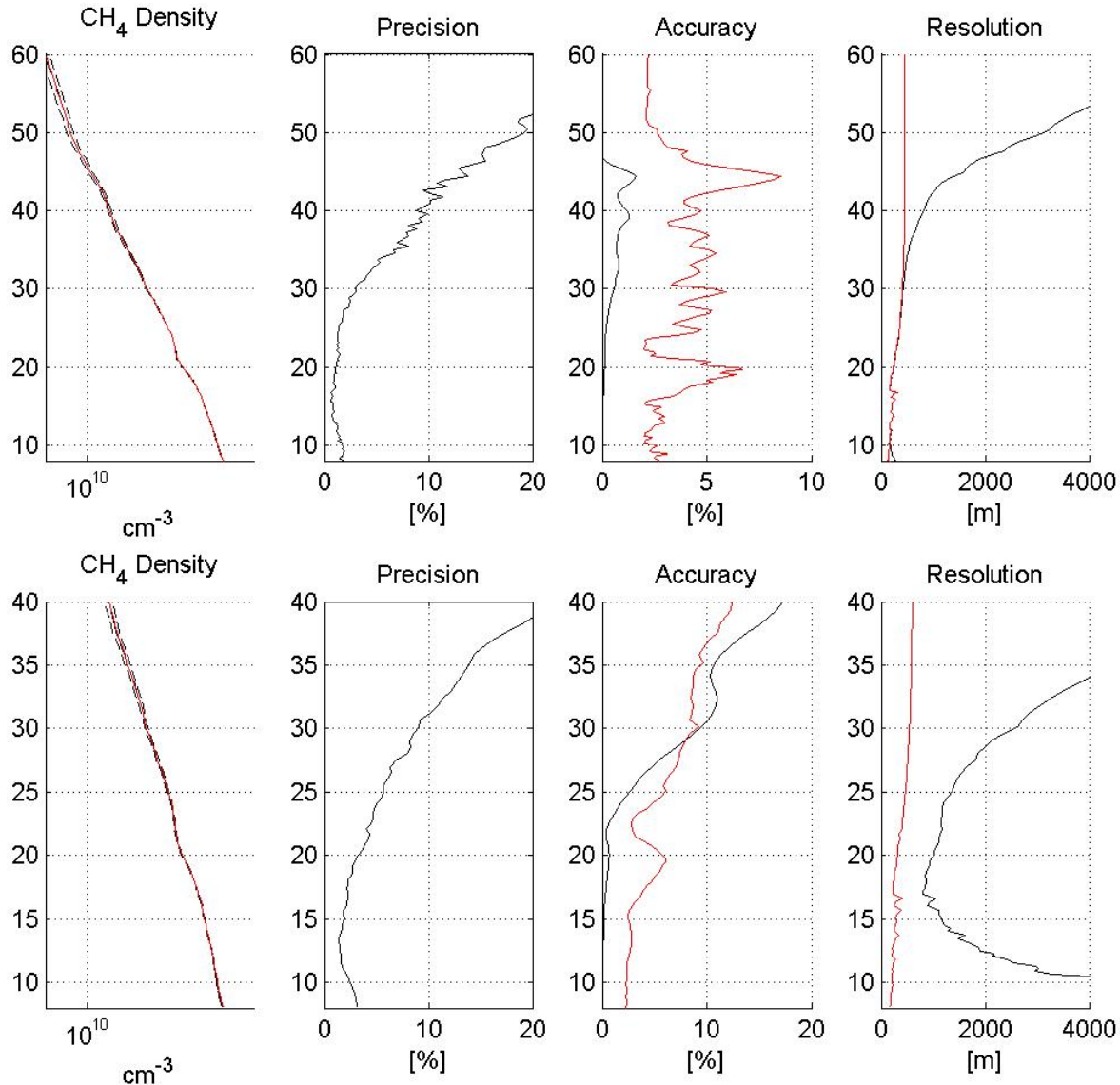
Stellar occultation:

- Sirius (αCMa)
- Magnitude: -1.44
- Temp: 11000 K

*Main error source
(troposphere): pointing
error.*

Occultation results: methane

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Solar occultation

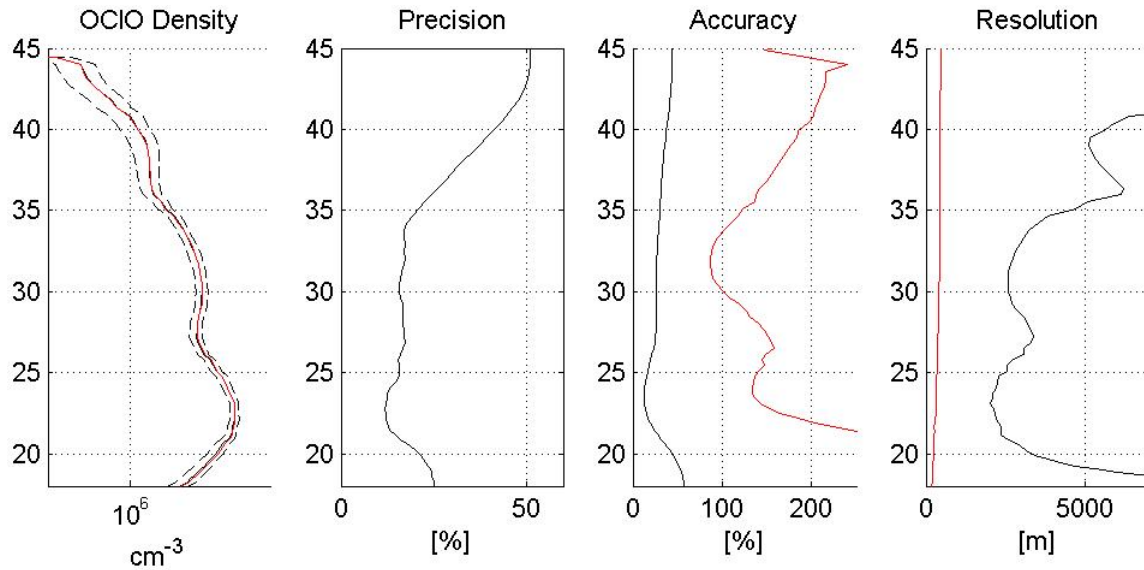
Stellar occultation:

- Sirius (αCMa)
- Magnitude: -1.44
- Temp: 11000 K

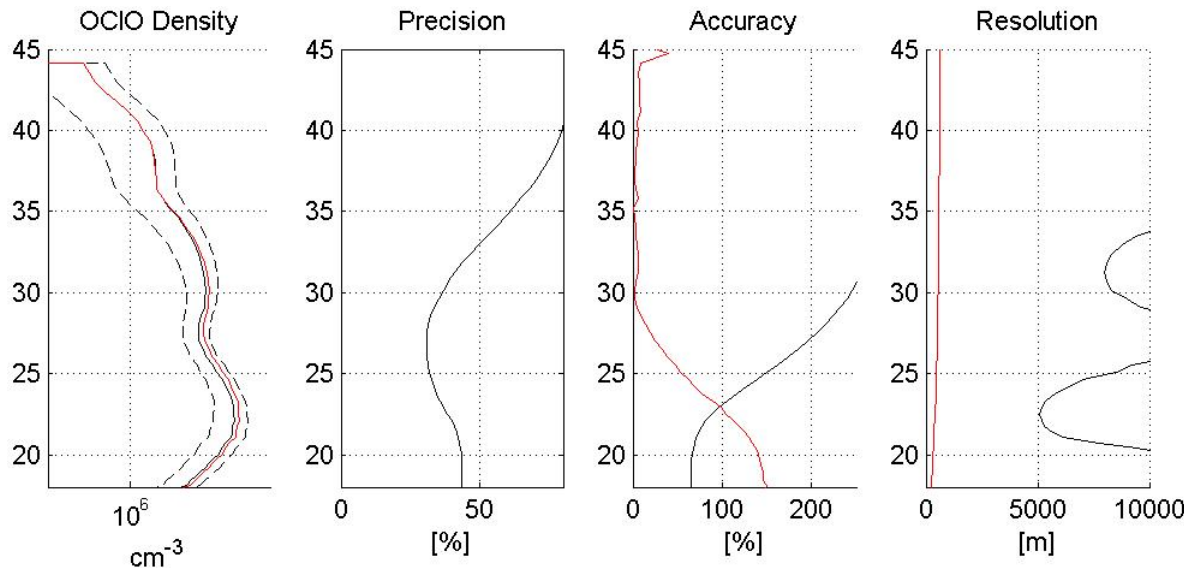
*Main error source:
pointing error*

Occultation results: OCIO

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Solar occultation



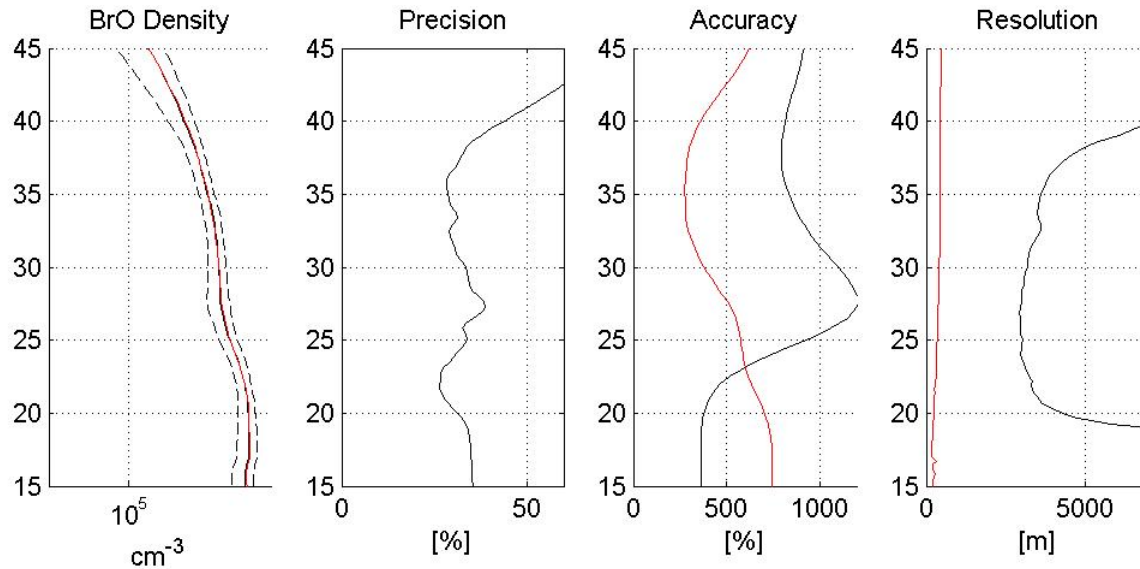
Stellar occultation:

- Sirius (α CMa)
- Magnitude: -1.44
- Temp: 11000 K

Pointing error as well as wavelength assignment error are very large

Occultation results: BrO (solar occ. only)

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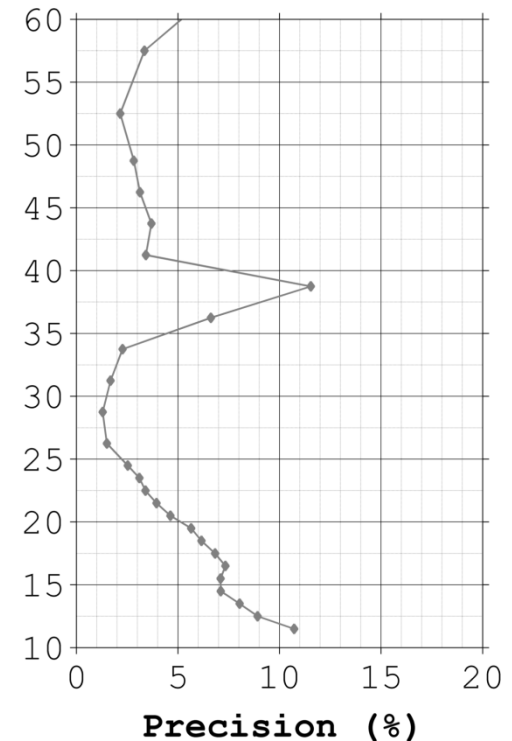
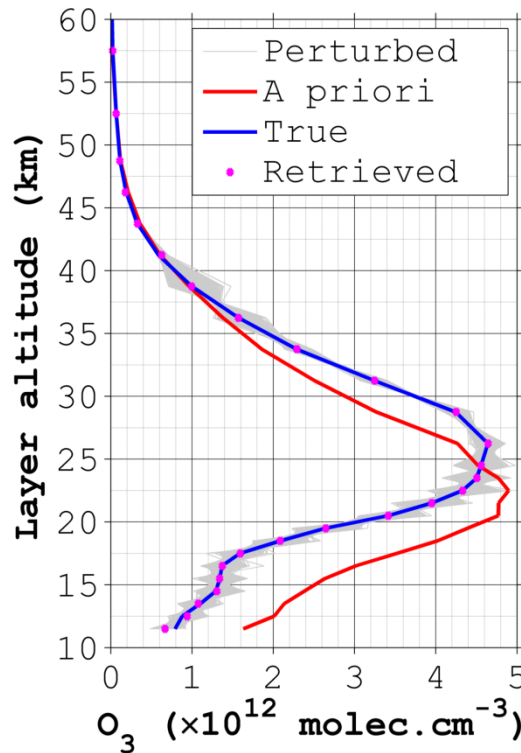
Solar occultation

*Wrong wavelength
assignment induces
huge errors*

ALTIUS bright limb O3

Error budget for O3 after random perturbation of the signal within the SNR requirements.

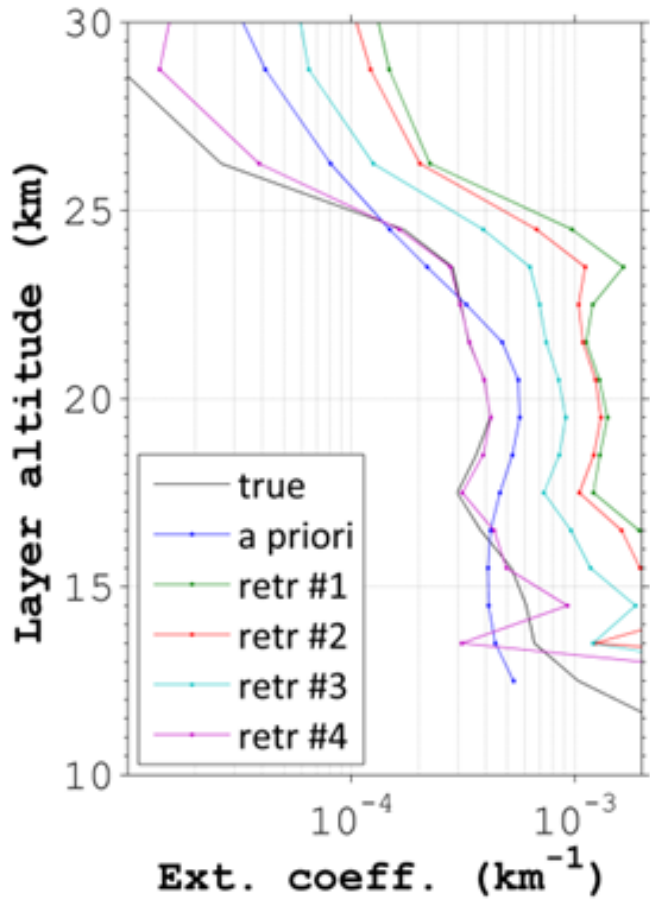
Total error budget including contribution from pointing and spectral registration.



Error source	15km	20km	25km	30km	35km	40km	45km	50km	55km
SNR	7%	5%	2%	1.5%	2.5%	3%	3.5%	2.5%	2.5%
Pointing	2%	4%	1%	1%	2%	2.5%	2.5%	3%	3%
Spectral	0%	0%	0%	0%	0%	0.2%	0.4%	0.5%	0.6%
Total	7.3%	6.4%	2.2%	1.8%	3.2%	3.9%	4.3%	3.9%	4%

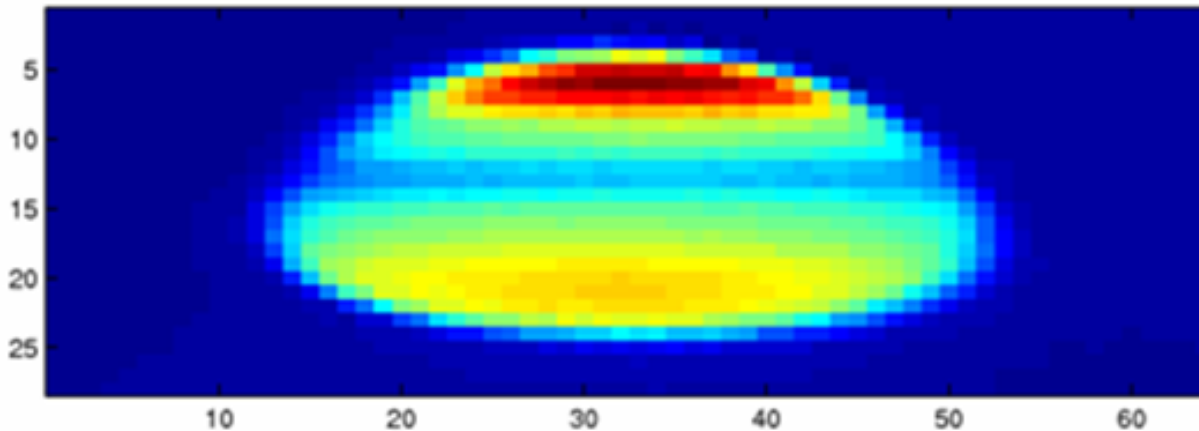
ALTIUS bright limb aerosol extinction

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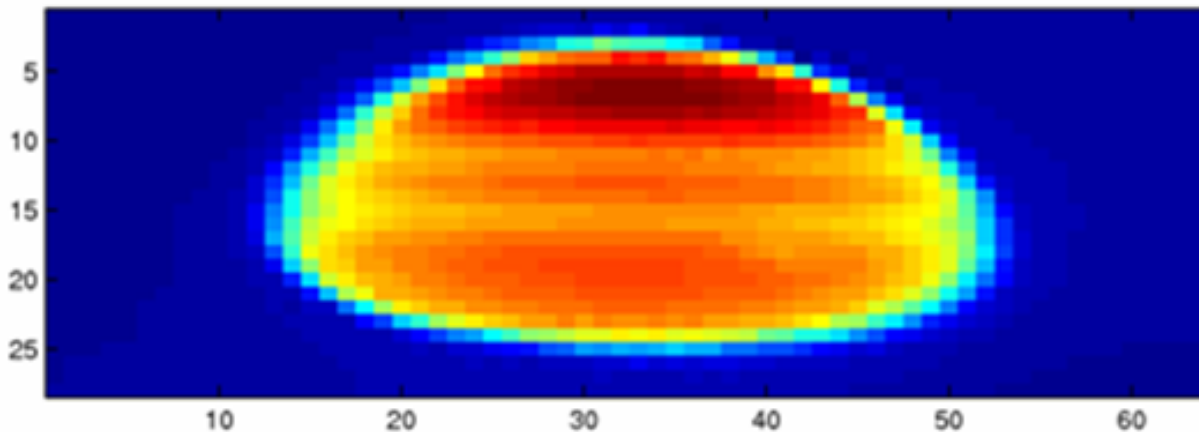


	Error Type	15km	20km	25km	30km
SNR	random	0.1%	0.1%	0.1%	0.1%
Pointing	bias	3%	2%	10%	10%
Spectral	bias	0%	0%	0%	0%
Total		3%	2%	10%	10%

Major advantage of imagers: you know what you are looking at.



- Canadian ACE instrument: sunset images at 1 micrometer.
- Single and double-layer PSC
- Every cloud type has its own distinct morphology (cirrus, PSCs, PMCs, tropical anvil clouds)



Bonus: AOTF-based imagers do work!

ESA-EO **AROMAT-2** campaign in Romania in Aug 2015.

Goal: intercompare different instruments around big pollution spots in view of TROPOMI ground-based validation.

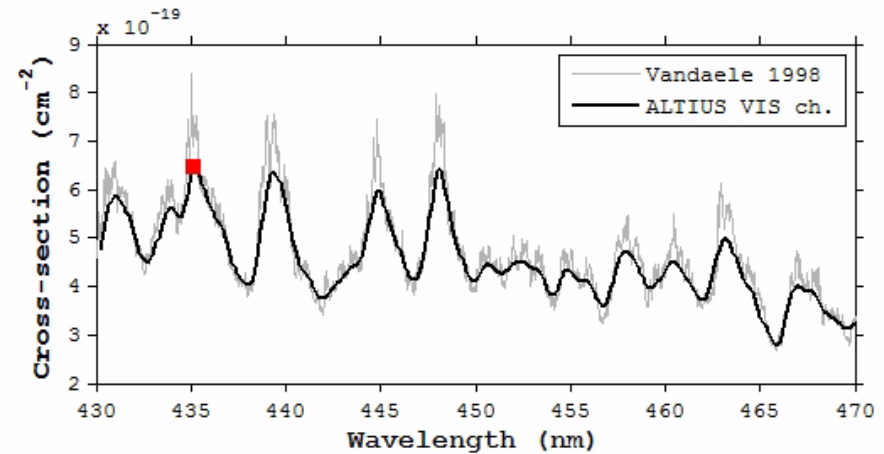
ALTIUS VIS channel breadboard was operated to retrieve NO₂ SCD from the plume of the coal-burning power plant of Turceni.



Bonus: AOTF-based imagers do work!

NO₂ retrieved from the cross-section features by comparing doublets of wavelengths.

$\lambda = 435.1\text{nm}$, 24-Aug-15 16:16:38.500



Bonus: AOTF-based imagers do work!

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